



# Looking forward to new Physics with FASER

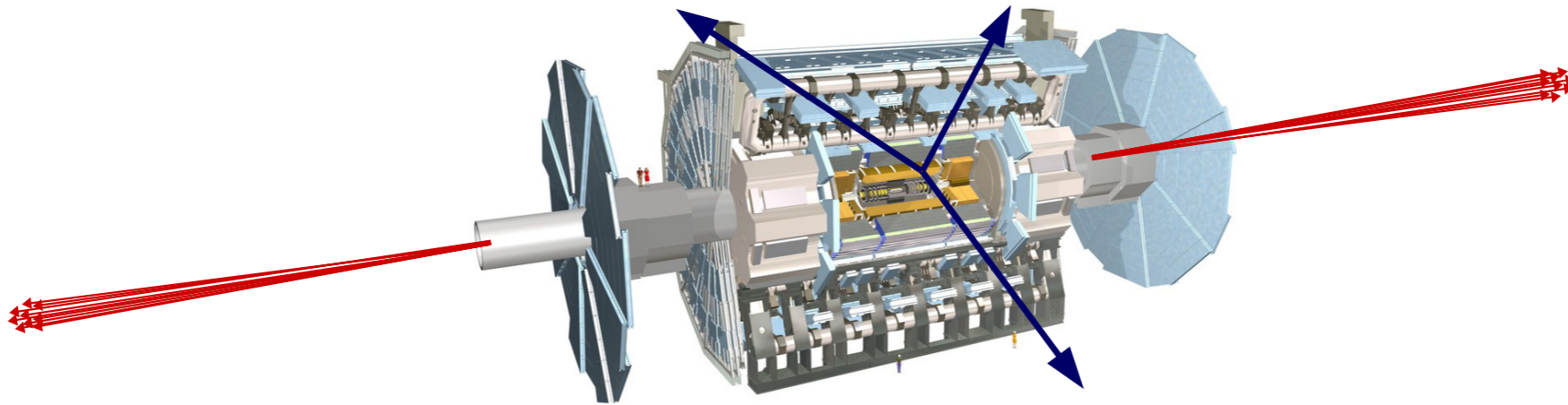
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October 1st 16th 2020  
Oklahoma State HEP seminar



# Motivation and Idea

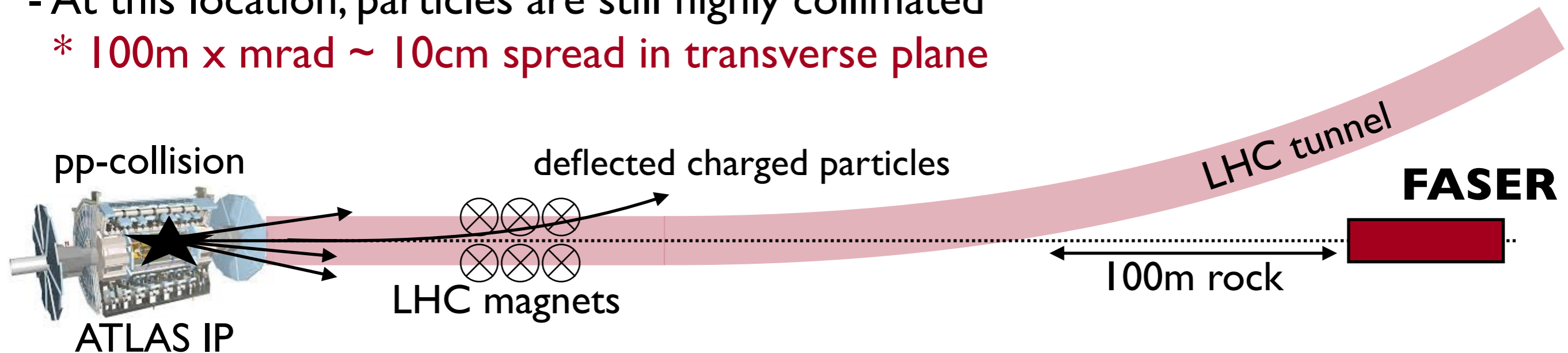
- LHC searches/experiments focus on **central region**, which is motivated by heavy, strongly interacting particles
  - \* small rates:  $\sigma \sim \text{fb} - \text{pb}$  or  $N_H \sim 10^7$  at  $\mathcal{L} = 300 \text{ fb}^{-1}$
  - \* high  $p_T$ , produced  $\sim$  isotropical



- For light and weakly interacting particles, this may be completely misguided
  - \* light: we can produce them in  $\pi, K, D, B$  decays
  - \* weakly-interacting: need extremely large SM event rate to see them
- We should go where the pions are: **forward region** along the beam line
  - \* enormous event rates:  $\sigma_{\text{inel}} \sim 100 \text{ mb}$  or  $N_\pi \sim 10^{17}$  at  $\mathcal{L} = 300 \text{ fb}^{-1}$
  - \* highly energetic beam remnants:  $E \sim \text{TeV}$
  - \* low  $p_T \sim \Lambda_{\text{QCD}} \rightarrow$  particles are collimated  $\theta \sim \Lambda_{\text{QCD}}/E \sim \text{mrad}$

# Motivation and Idea

- We can't place a reasonably-sized detector on the beam line near the IP
  - \* blocks the proton beams, subject to large radiation
- However, weakly-interacting particles do not interact with matter
  - place detector few 100m away along the “collision axis” after beam curves
  - \* LHC infrastructure acts and rock act as shielding
- At this location, particles are still highly collimated
  - \*  $100\text{m} \times \text{mrad} \sim 10\text{cm}$  spread in transverse plane

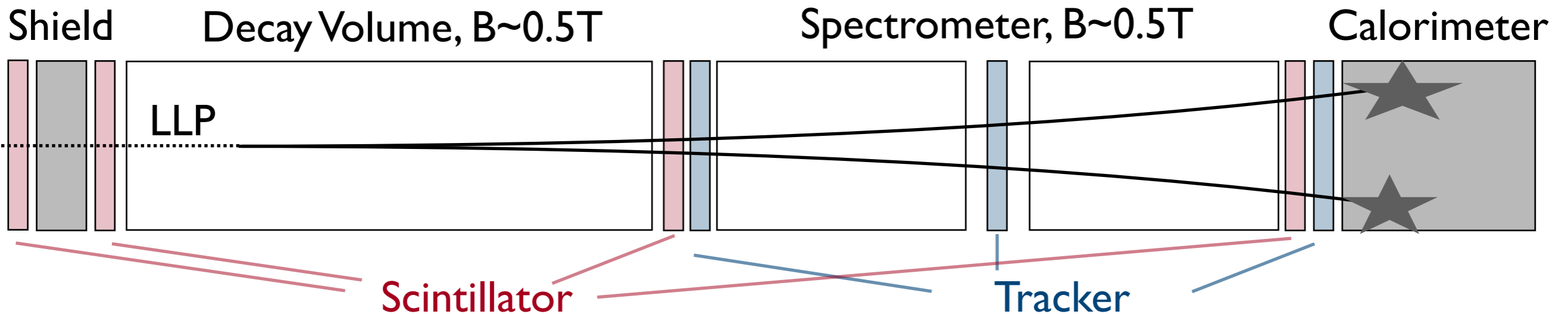


- This motivates small, fast and cheap inexpensive detector  
**FASER: ForwArd Search ExpeRiment** at the LHC
- Applications for light long-lived particles searches and neutrinos

# Motivation and Idea

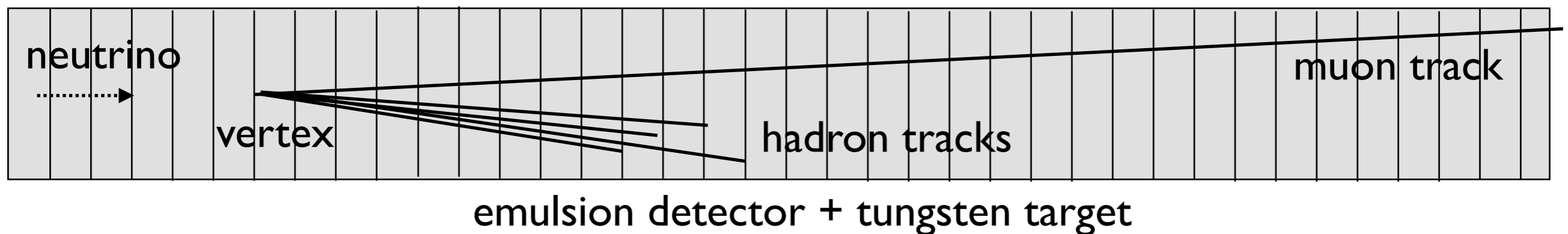
## Long-Lived Particles Searches

search for appearance of high-energy particles from LLP decay in empty decay volume



## Neutrino Measurements

search for neutrino interaction vertex in dense material using emulsion detector



# Outline

## **The FASER Experiment**

Location / Detector / Environment / Timeline

## **Long-Lived Particle Searches**

Dark Photons

## **Neutrino Physics**

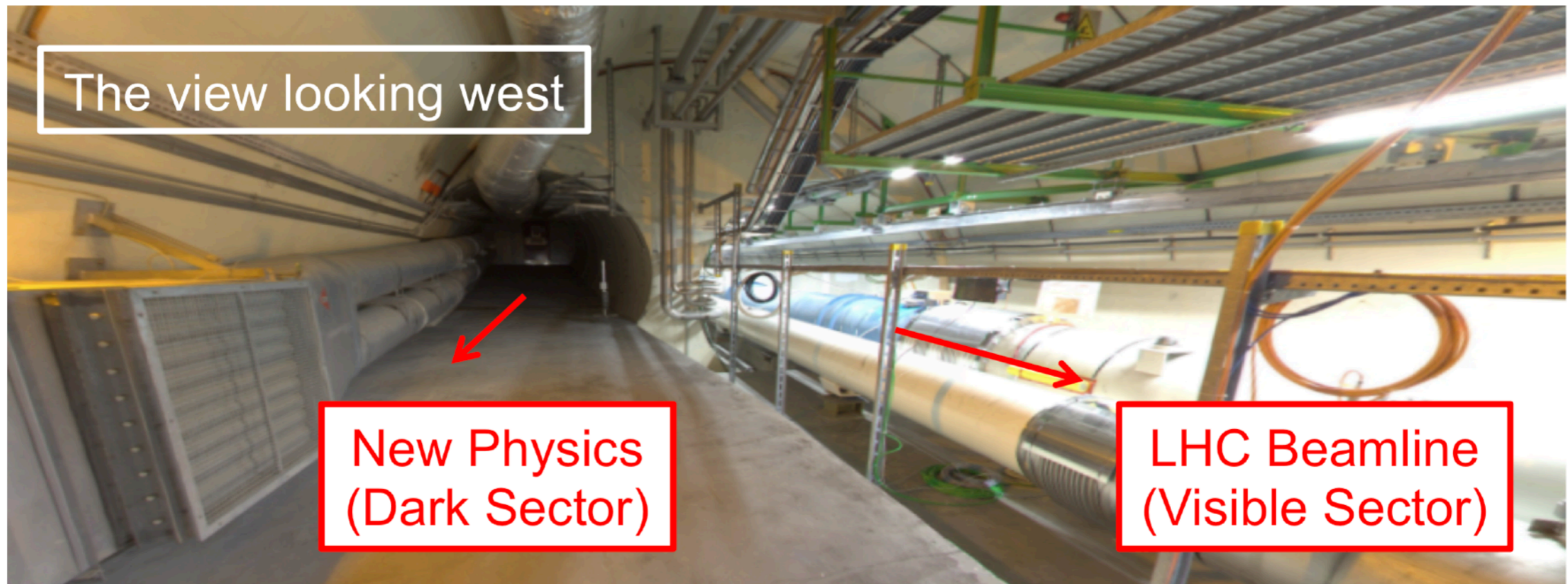
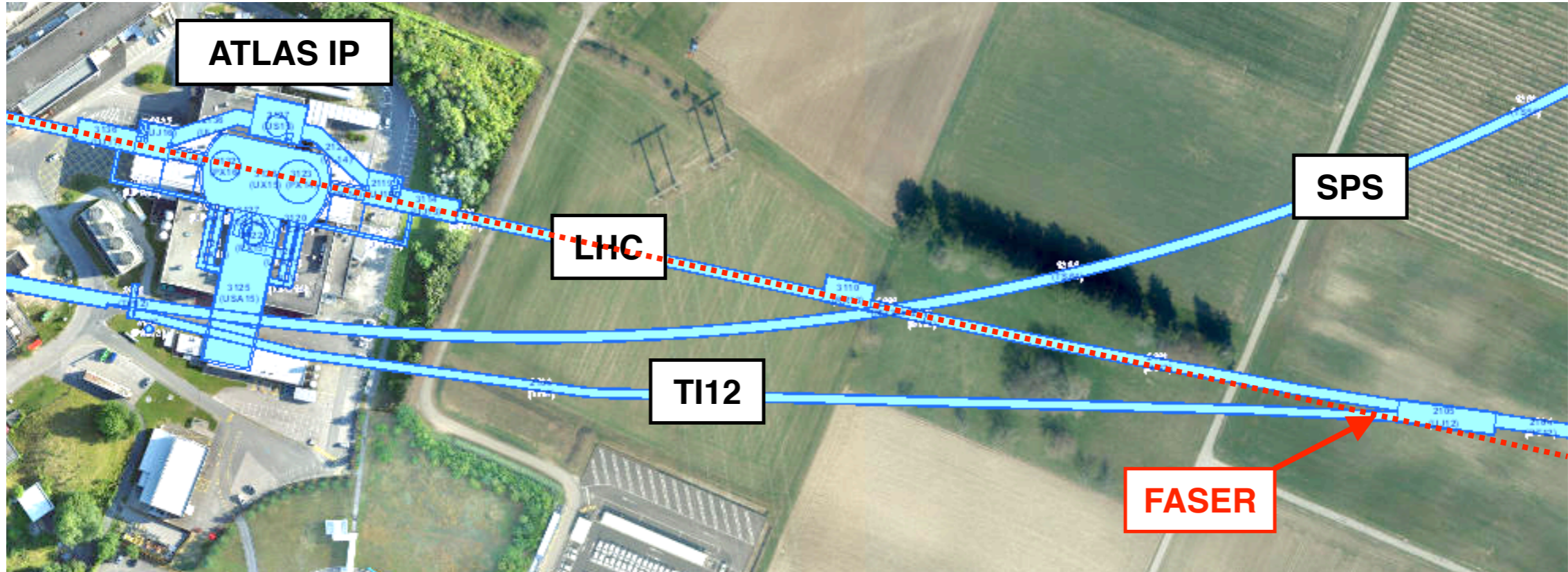
FASER $\nu$  / Neutrino Measurements

## **Summary and Conclusion**

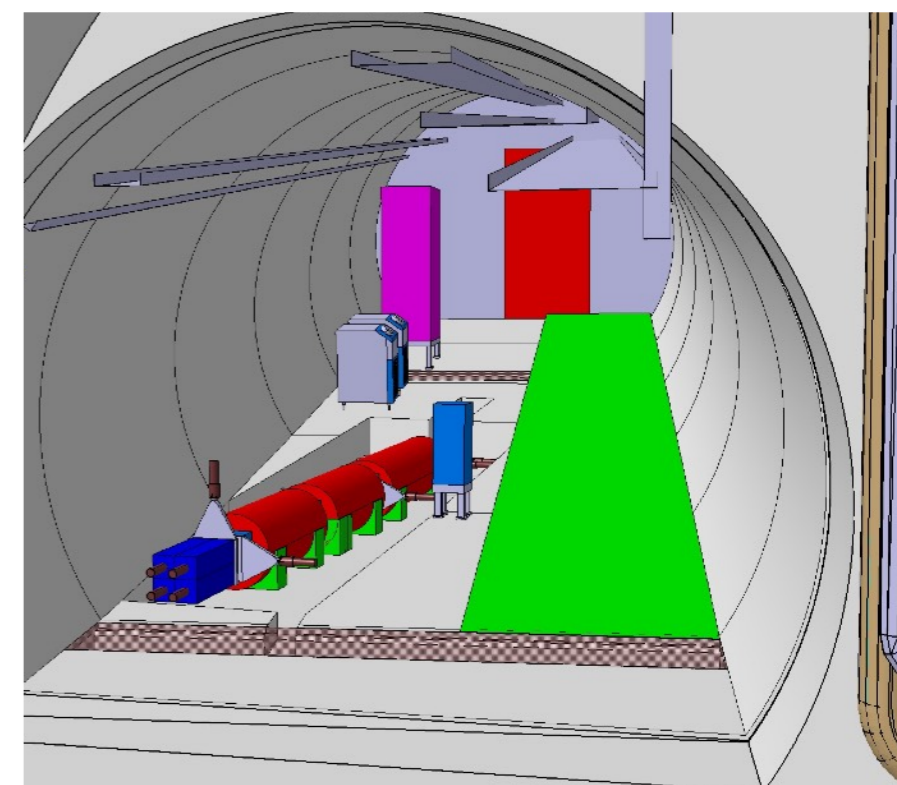


## **The FASER experiment**

# Location



# FASER Location



FASER location is being prepared





# FASER Detector

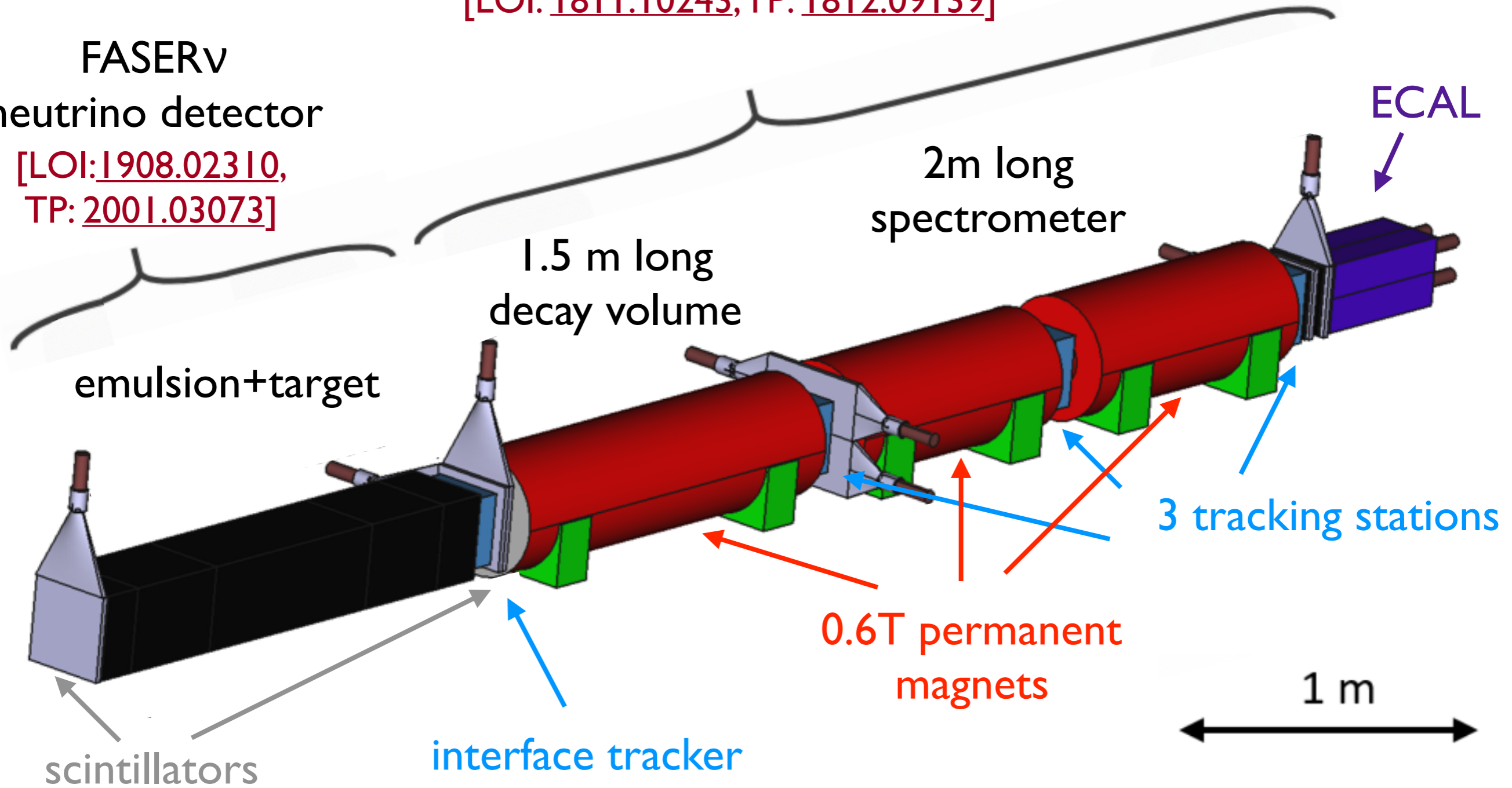
FASER main detector

[LOI: [1811.10243](#), TP: [1812.09139](#)]

FASER $\nu$

neutrino detector

[LOI: [1908.02310](#),  
TP: [2001.03073](#)]



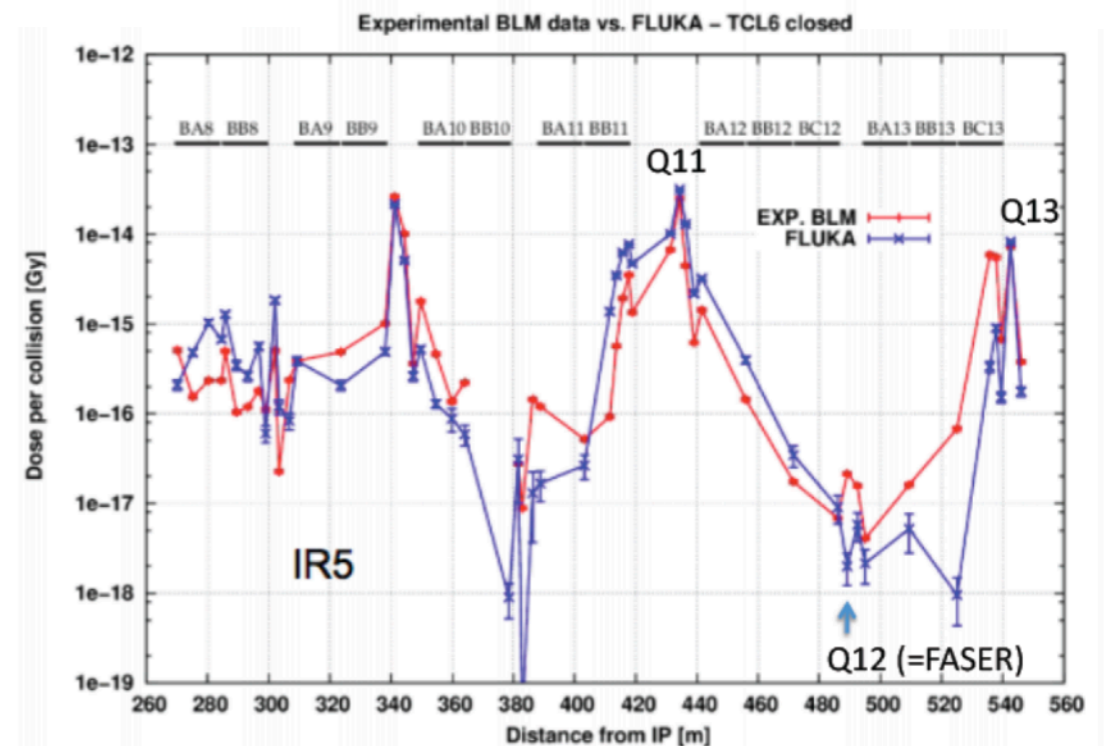
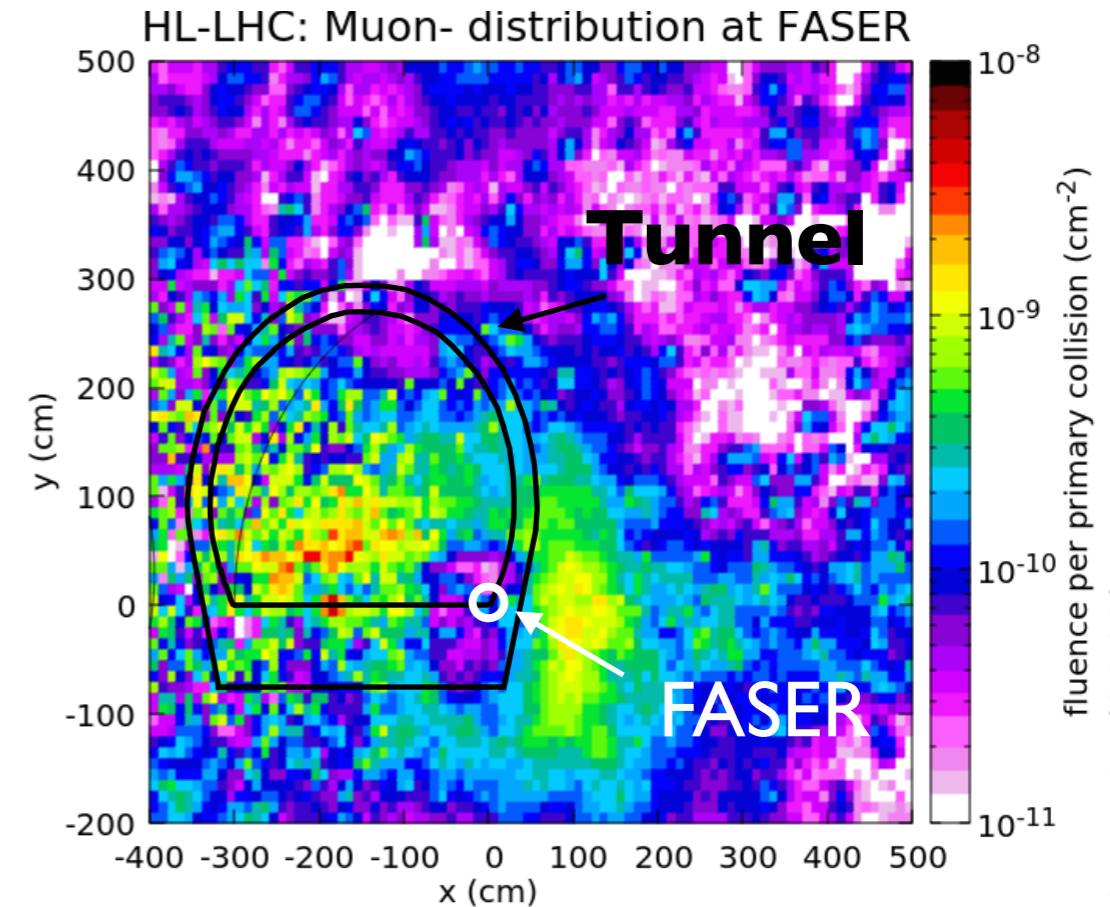
# FASER Environment

## General Considerations

- LLP/neutrino signals have TeV energies, come from ATLAS IP
  - shielding: natural (rock) and LHC infrastructure (magnets, absorbers)
  - only muons/neutrino can transport TeV energies through  $\sim 100\text{m}$  rock
- FASER's location is very quiet

## FLUKA simulation (by CERN STI group)

- FASER in remarkable quiet spot
  - estimated muon flux from IP:  $2 \cdot 10^4 \text{ fb/cm}^2$
  - HE particles from beam-gas collisions and proton losses near FASER are negligible
  - other HE particles produced in muon radiative processes
- use scintillator to veto muons for LLP searches

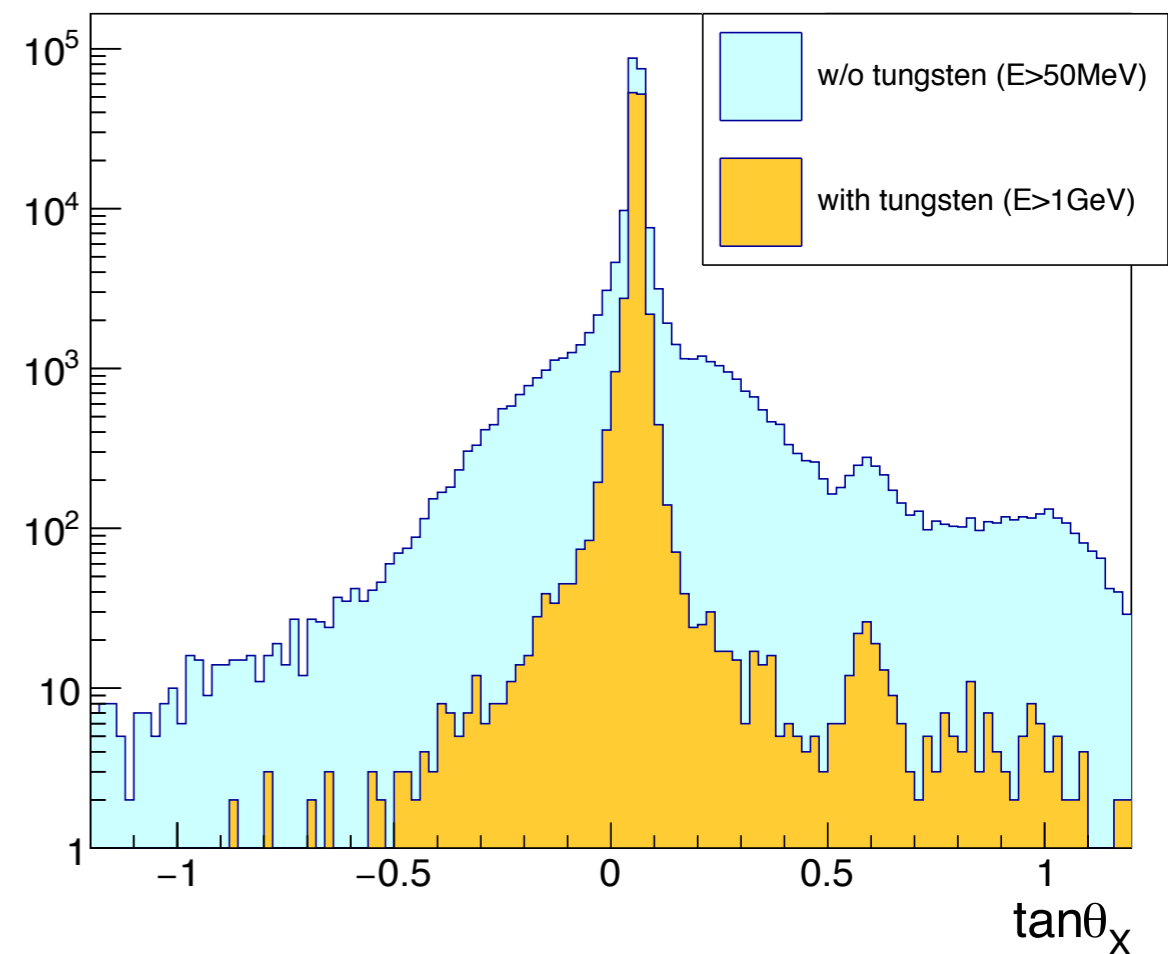


Sabate-Gilarte, Cerutti, Tsinganis (2018)

# FASER Environment

## In-Situ Measurements

- BatMon radiation monitor:
  - \* low-energy radiation levels are promisingly low
- emulsion detector installed in 2018 in both T118 and T112
  - \* consistent with FLUKA simulations
  - \* more data analysis on-going





## Long Lived Particles Searches at FASER

Idea: [1708.09389](#)

LOI: [1811.10243](#)

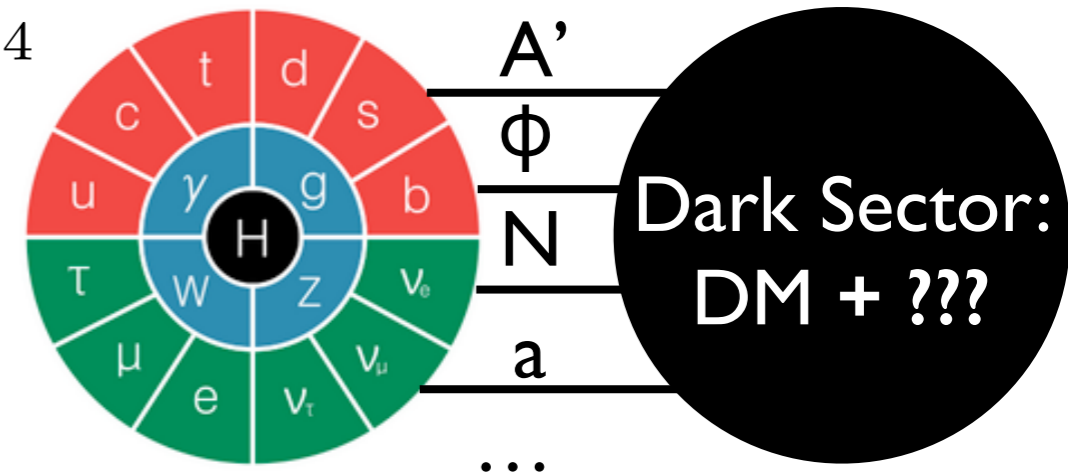
TP: [1812.09139](#)

Physics: [1811.12522](#)

# Long Lived Particles at FASER

## Motivation

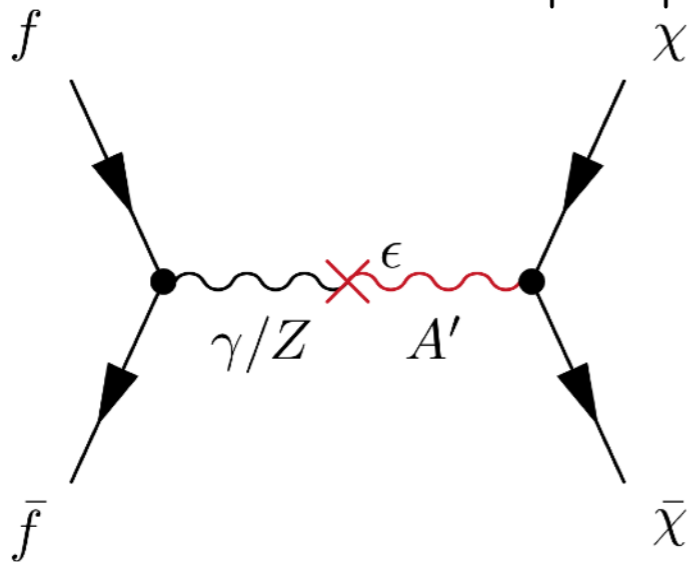
- Dark matter solid evidence for new particles
  - \* thermal freeze out:  $\Omega_{DM} \sim 1/\langle\sigma v\rangle \sim m^2/g^4$
  - \* WIMP miracle:  $m \sim m_{weak}, g \sim g_{weak}$
  - \* light DM ( $m \sim \text{GeV}$ ) requires new mediators
    - $m < m_{weak}, g < g_{weak}$
    - light weakly coupled particles
- Anomalies: muon  $g-2$ , Be-8, KOTO, ...



## Prominent examples

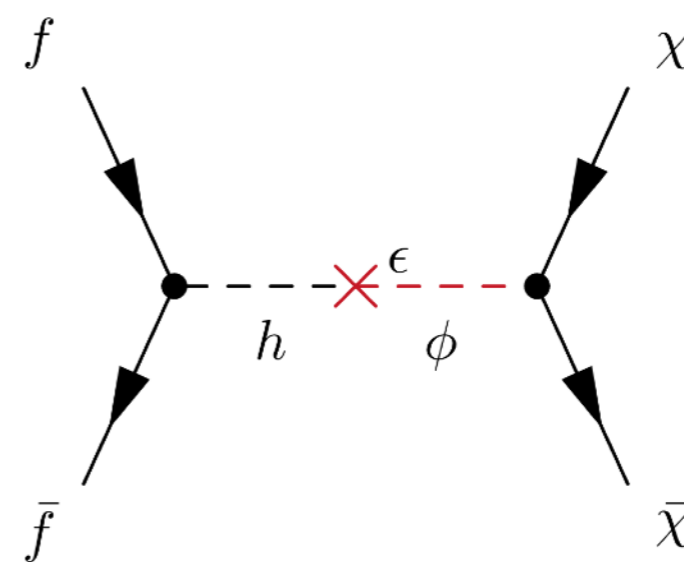
Dark Photon Portal:  $\epsilon F^{\mu\nu} F'_{\mu\nu}$

Dark Higgs Portal:  $\epsilon |H|^2 \phi^2$



Neutrino Portal:  $y L H N$

Axion Portal:  $g a F^{\mu\nu} \tilde{F}_{\mu\nu}$



# Long Lived Particles at FASER

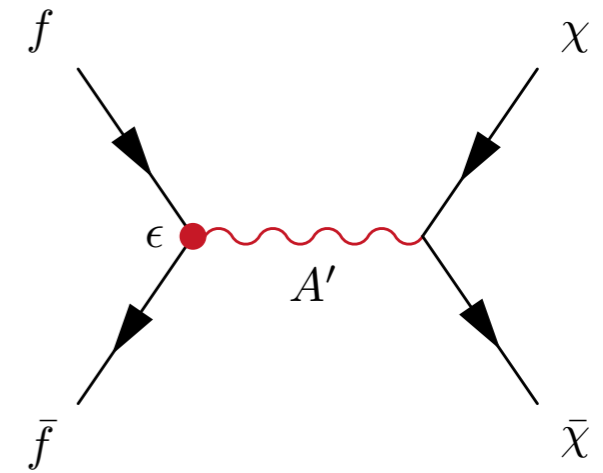
## Motivation

- dark matter solid evidence for new particles
- light DM ( $m \sim \text{GeV}$ ) requires new mediators  
→ light weakly coupled particles

## Dark Photons

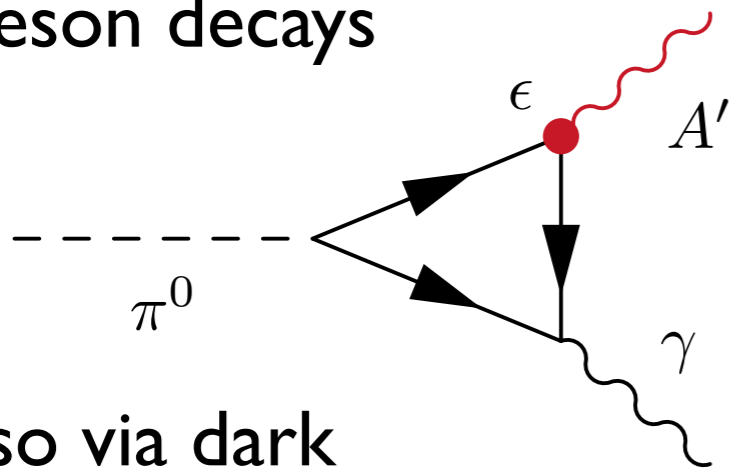
- similar to the SM photon but
  - \* massive, with mass  $m_{A'}$
  - \* couplings to SM particles suppressed by  $\epsilon$

$$\mathcal{L} = \frac{1}{2} m_{A'}^2 A'_\mu A'^\mu + \sum \bar{f} (i \not{\partial} - \epsilon e q_f A') f$$



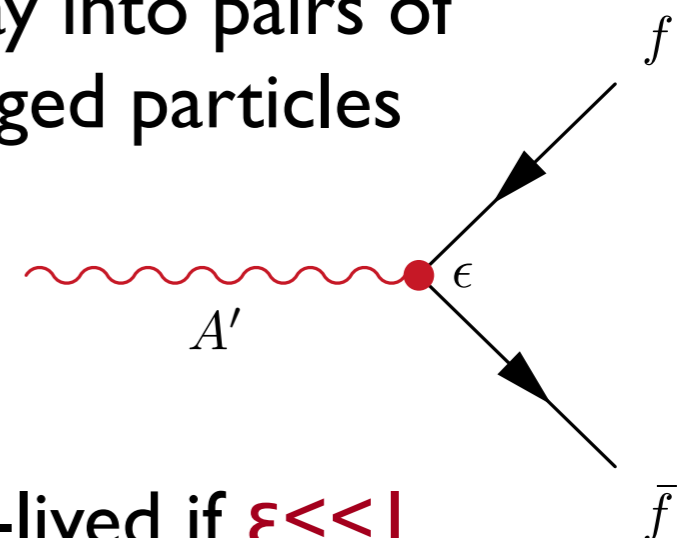
## Dark Photons at FASER

- produced for example in meson decays



- also via dark Bremsstrahlung at large  $m_{A'}$

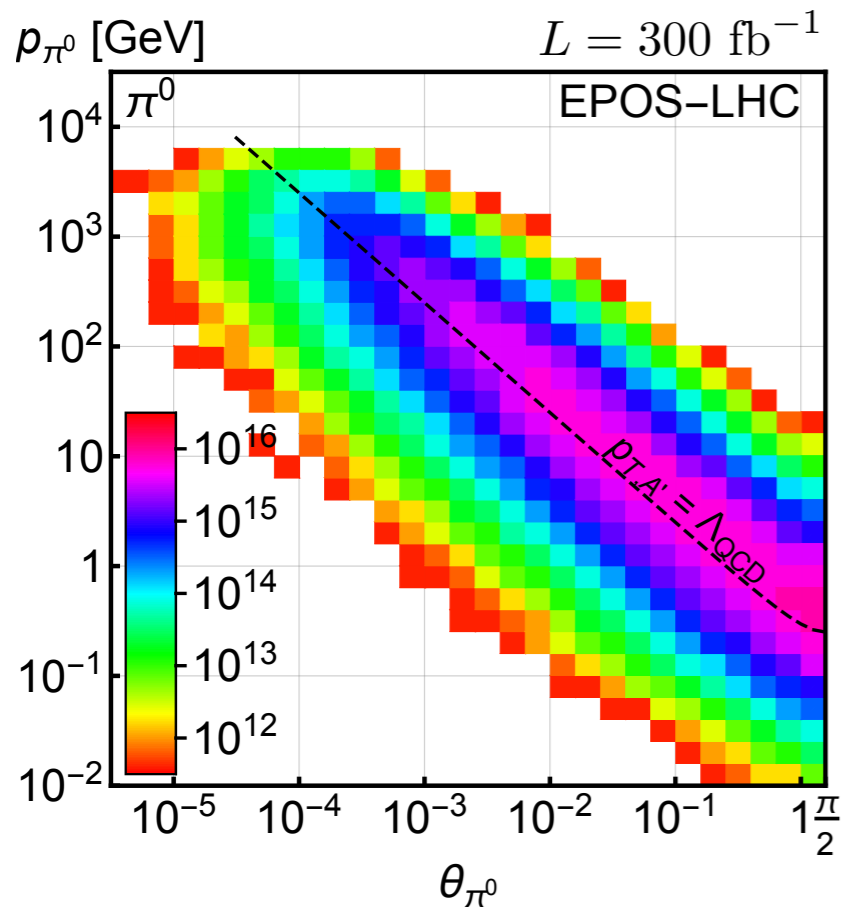
- decay into pairs of charged particles



- long-lived if  $\epsilon \ll 1$

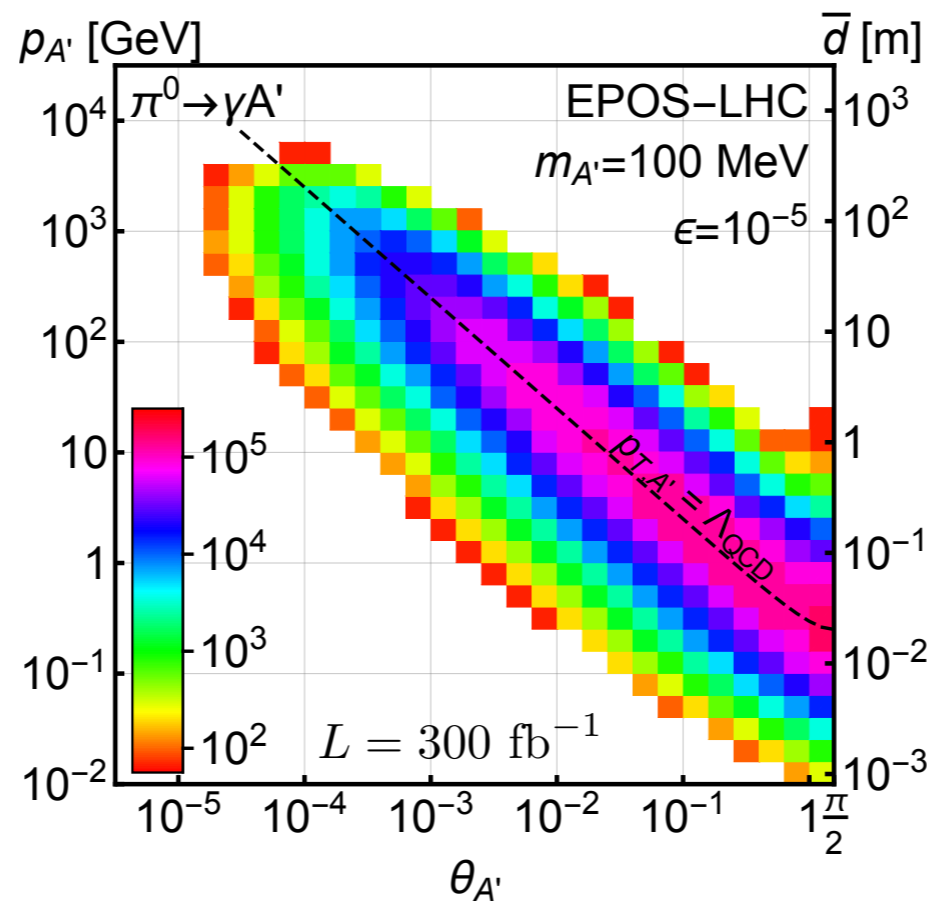
# Long Lived Particles at FASER

## Pions at IP



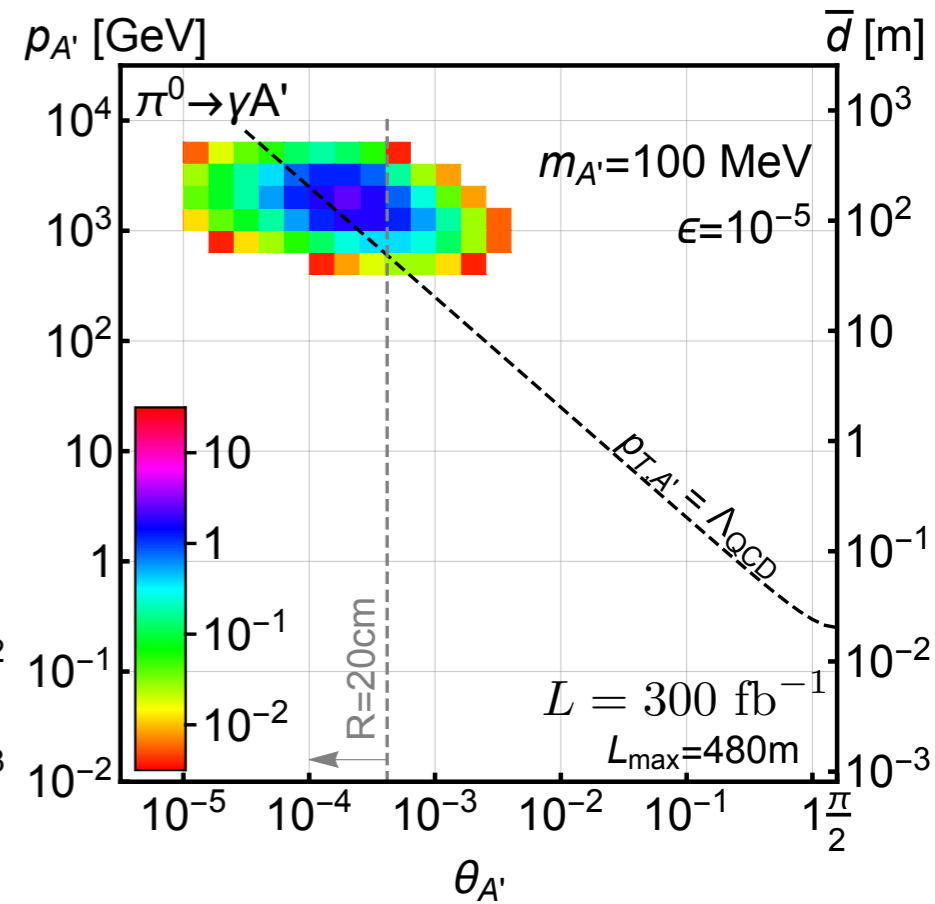
- dedicated hadronic interaction models, grounded on LHC data
- production peaks at  $p_T \sim \Lambda_{QCD}$
- enormous event rates  $N \sim 10^{15}$  per bin

## A' at IP



- production peaks at  $p_T \sim \Lambda_{QCD}$
- rates highly suppressed by  $\epsilon^2 \sim 10^{-10}$
- still rates  $N \sim 10^5$  per bin: LHC could be dark a photon factory

## A' decay at FASER



- only highly boosted  $\sim \text{TeV } A'$  arrive at FASER
- rates suppressed by decay requirements
- still rates  $N \sim 100$  signal events within 20cm of beam collision axis

# Long Lived Particles at FASER

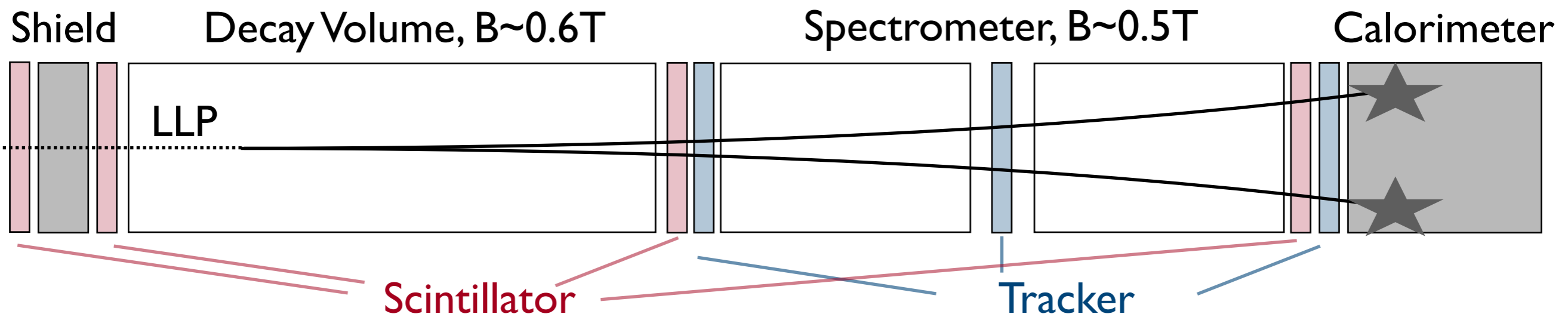
$pp \rightarrow \text{LLP} + X$ ,      LLP travels  $\sim 480\text{m}$ ,       $\text{LLP} \rightarrow \text{charged tracks} + X$

## Signal is striking

- two opposite-sign, high energy ( $E > 500 \text{ GeV}$ ) charged particles
- originate from a common vertex in a small, empty decay volume
- point back to the IP through 90 m of rock

## Background considerations

- cosmic rays and neutrino interactions (different kinematics) not a problem
- HE muon-associated radiative events are leading BG if muon is not vetoed
- incoming muons can be identified using scintillators  
→ reduce BG to negligible levels



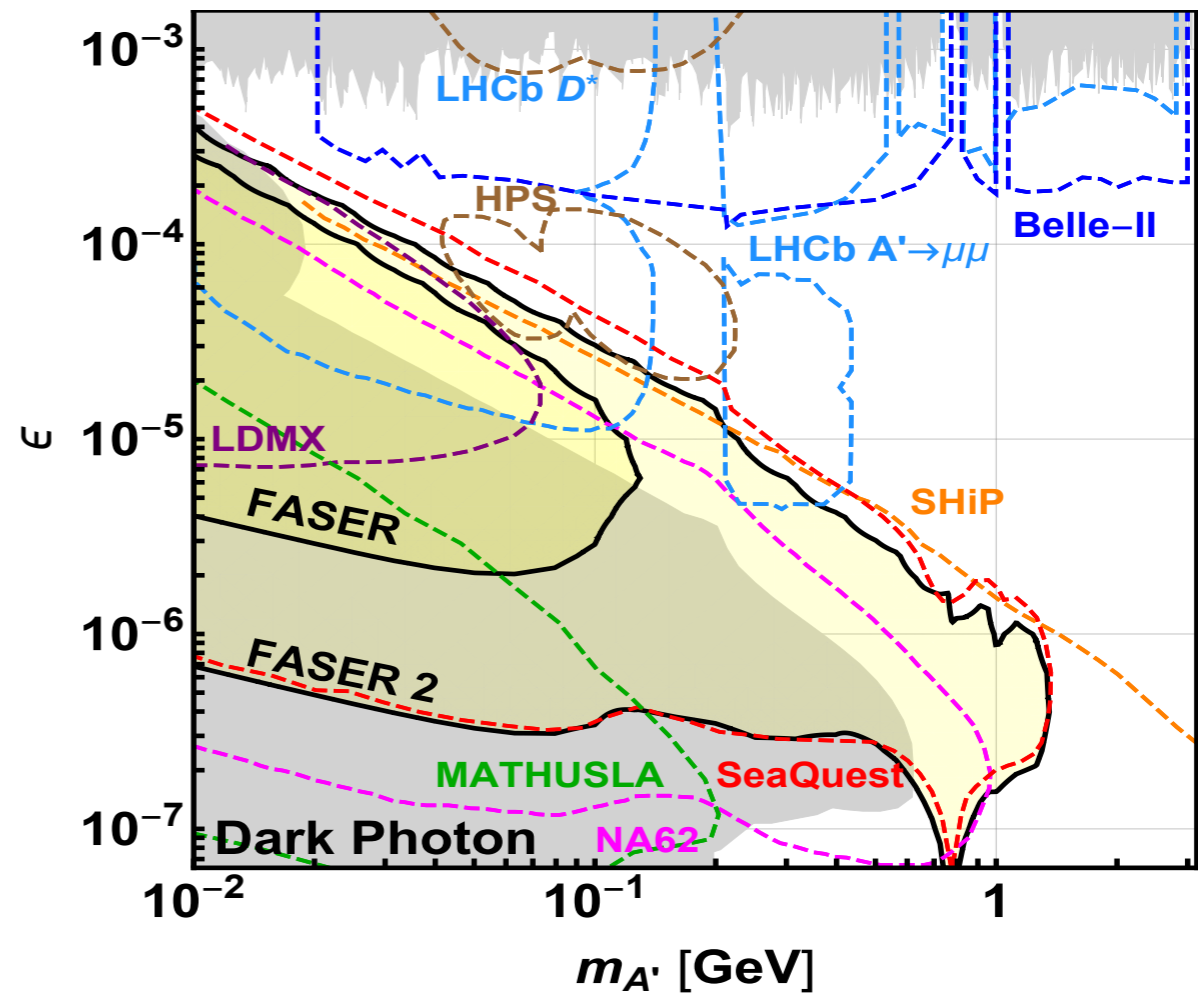
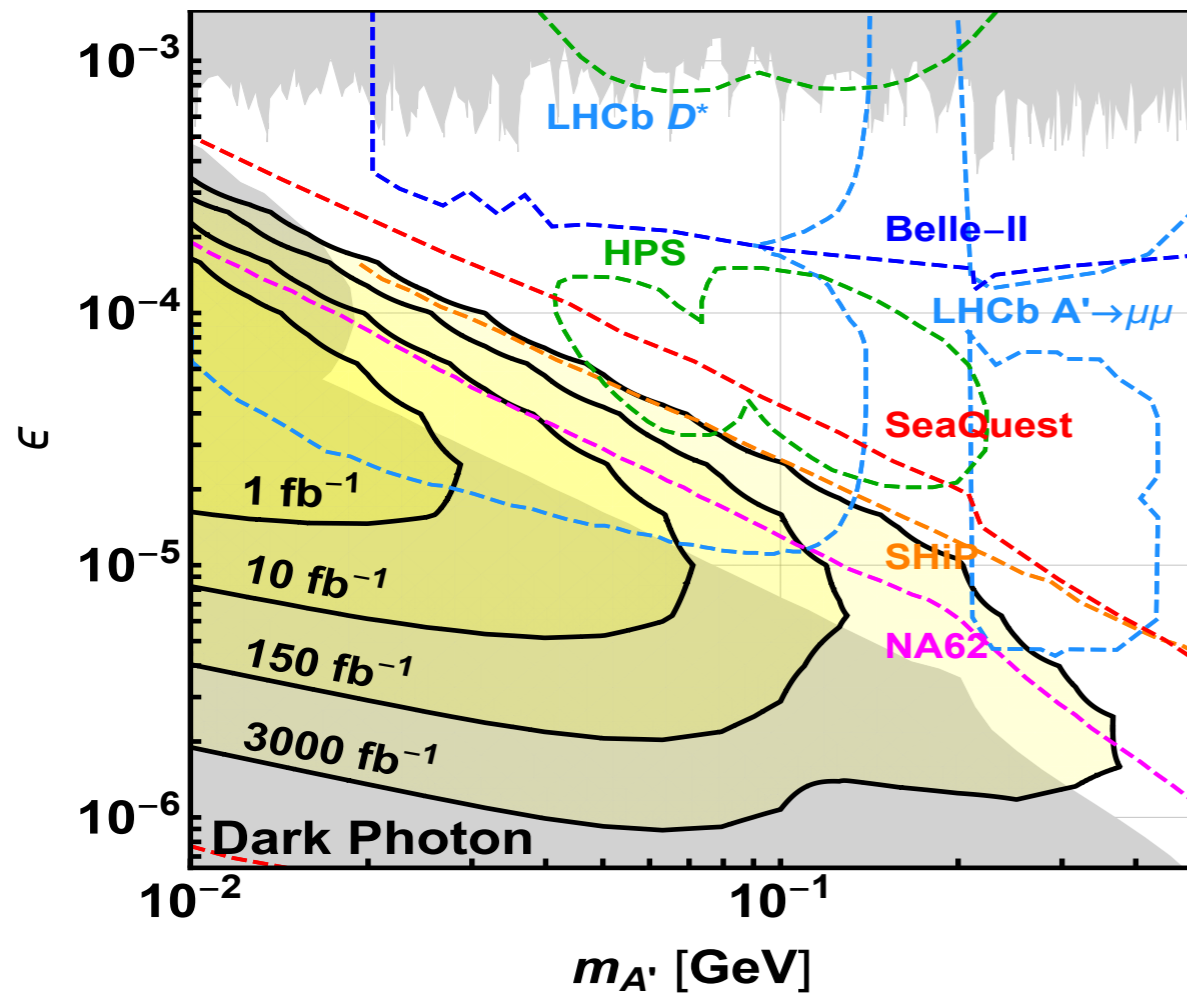


# Long Lived Particles at FASER

## Dark Photon Reach

- consider all production channels
- include  $A' \rightarrow ee, \mu\mu, \pi^+\pi^-$
- require  $N=3$  events

- two detector benchmarks
- \* FASER:  $R=10\text{cm}, L=1.5\text{m}, \text{Run3}$
- \* FASER 2:  $R=1\text{m}, L=5\text{m}, \text{HL-LHC}$

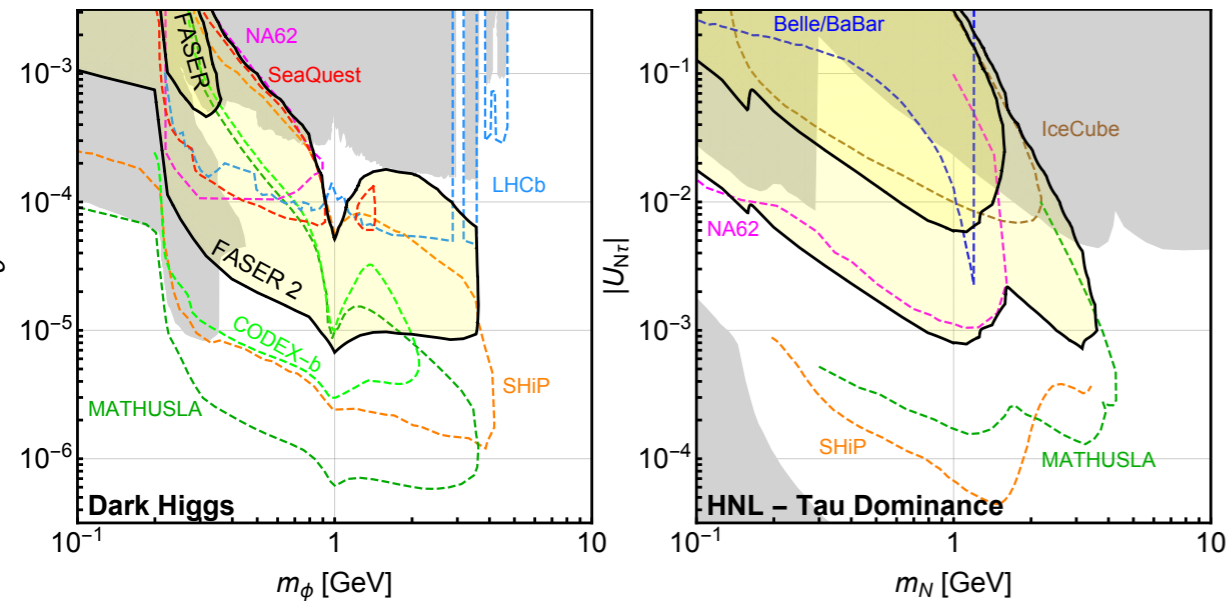


- reach limited by decay length (high  $\epsilon$ ) and production rate (low  $\epsilon$ )
- new parameter space probed with just  $1 \text{ fb}^{-1}$  in 2021
- FASER 2 significantly improves reach at higher masses

# Long Lived Particles at FASER

## Physics Sensitivity Studies

- FASER is sensitive to many more LLP models
- FASER Physics Case: [1811.12522](#)



- Many other studies related to FASER on FASER website: <https://faser.web.cern.ch/>

Benchmark Model	Label	Section	PBC	Refs	FASER	FASER 2
Dark Photons	V1	IV A	BC1	[7]	✓	✓
$B - L$ Gauge Bosons	V2	IV B	—	[30]	✓	✓
$L_i - L_j$ Gauge Bosons	V3	IV C	—	[30]	—	—
Dark Higgs Bosons	S1	V A	BC4	[26, 27]	—	✓
Dark Higgs Bosons with $hSS$	S2	V B	BC5	[26]	—	✓
HNLs with $e$	F1	VI	BC6	[28, 29]	—	✓
HNLs with $\mu$	F2	VI	BC7	[28, 29]	—	✓
HNLs with $\tau$	F3	VI	BC8	[28, 29]	✓	✓
ALPs with Photon	A1	VII A	BC9	[32]	✓	✓
ALPs with Fermion	A2	VII B	BC10	—	—	✓
ALPs with Gluon	A3	VII C	BC11	—	✓	✓
Dark Pseudoscalars	P1	VIII	—	[36]	—	✓

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- P. Bakhti, Y. Farzan, S. Pascoli, [Unravelling the richness of dark sector by FASERv](#)
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- C. Csaki, R. Tito D'Agnolo, M. Geller, A. Ismail, [Crunching Dilaton, Hidden Naturalness](#)
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## Neutrino Measurements at FASER

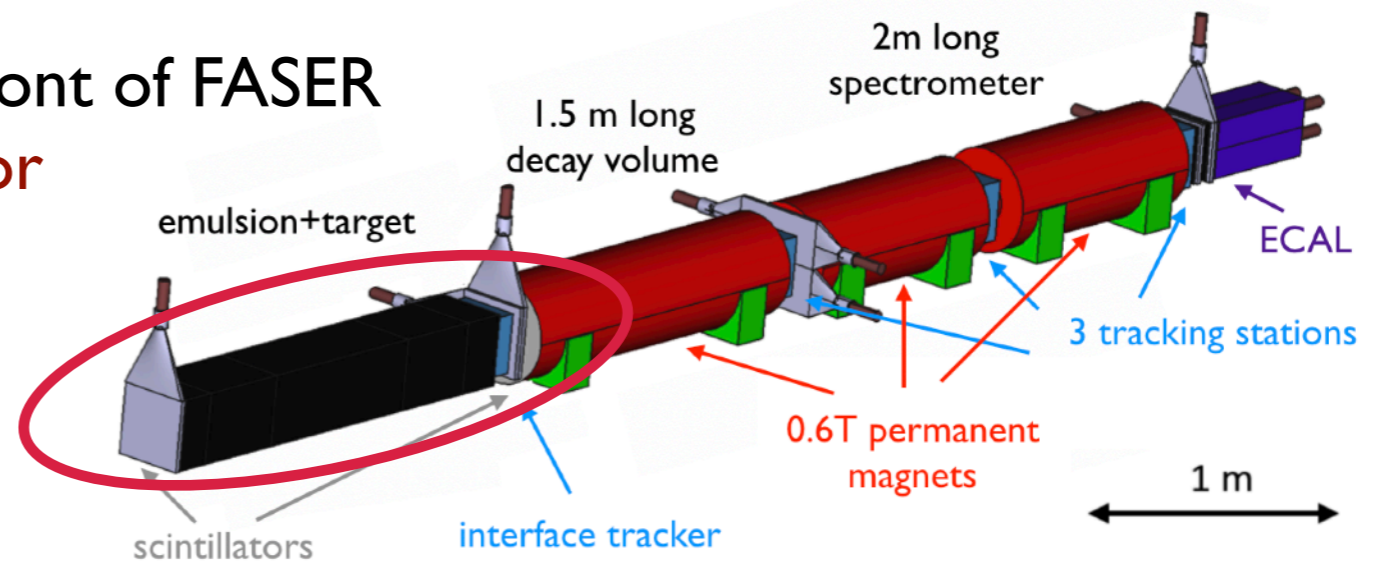
Idea/LOI: [1908.02310](#)  
TP: [2001.03073](#)

# FASER $\nu$ Motivation

- neutrinos detected from many sources, but not from colliders
- many neutrinos at LHC produced in  $\pi, K, D$  meson decay
  - ATLAS provides intense highly collimated neutrino beam towards FASER
  - \*  $\sim 10^{12}$  neutrino in LHC Run 3
  - \* highly collimated
  - \*  $E \sim \text{TeV}$

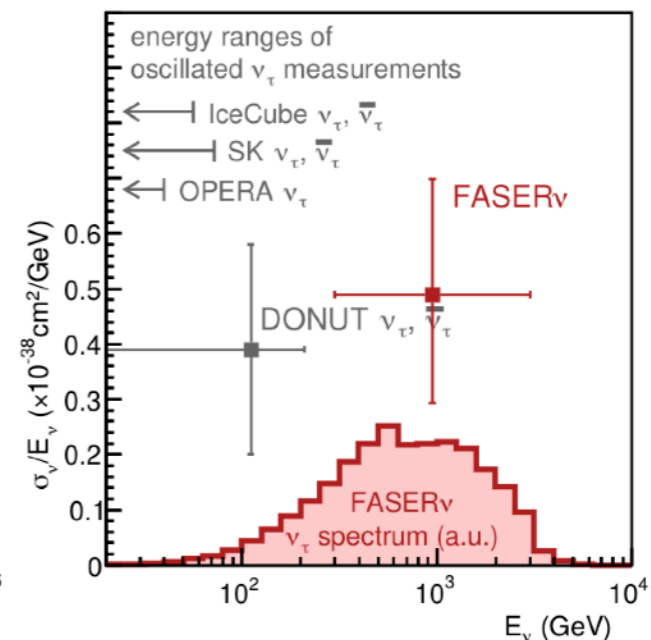
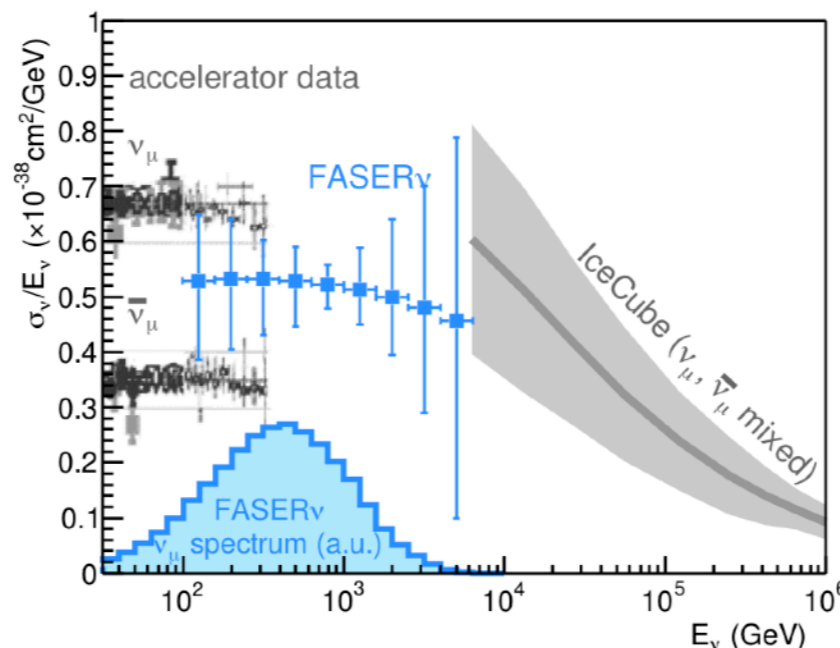
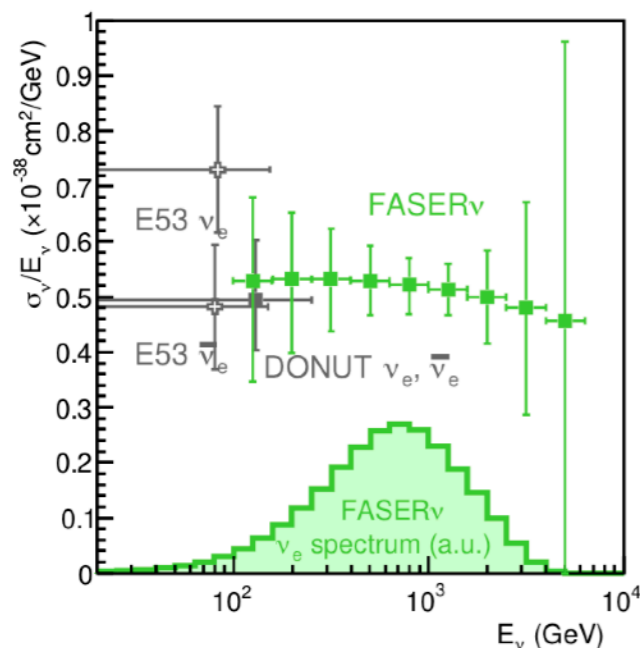
- dedicated FASER $\nu$  neutrino detector in front of FASER

- \* 25cm x 25cm x 1.3m emulsion detector
- \* tungsten target with 1.2 ton mass
- \*  $\sim 20000 \nu_\mu, \sim 2000 \nu_e, \sim 20 \nu_\tau$  during LHC Run 3



- neutrino physics at TeV energies

- \* probe unconstrained neutrino cross sections at TeV for all 3 flavors



# FASER $\nu$ Detector

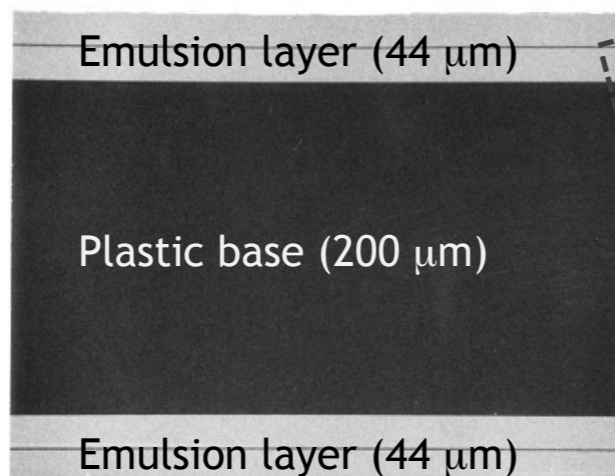
- emulsion detectors technology

\* used by many other neutrino experiments: CHORUS, DONUT, OPERA

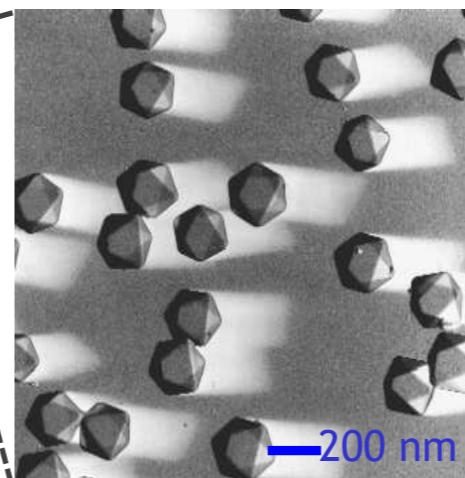
Emulsion film



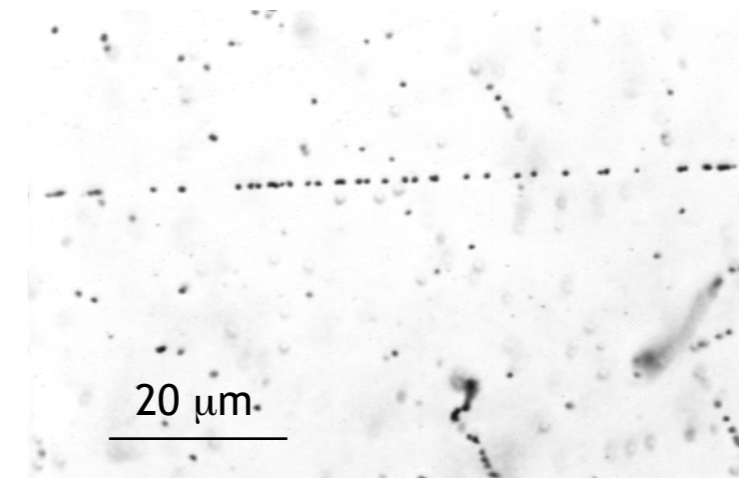
Cross-sectional view



AgBr crystal



Track in emulsion film

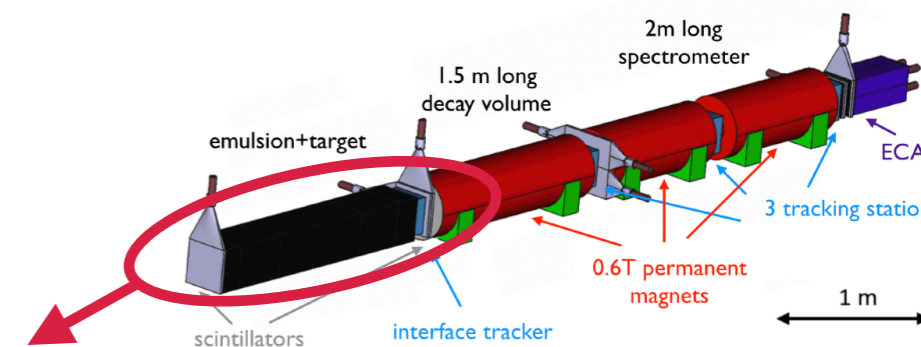
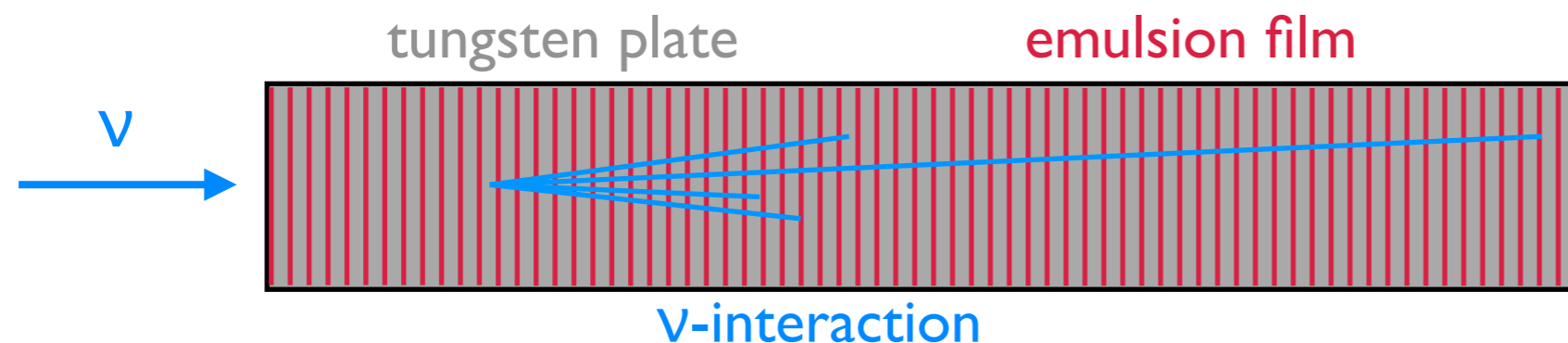


- FASER $\nu$ : 1000 emulsion films interleaved with 1mm tungsten plates

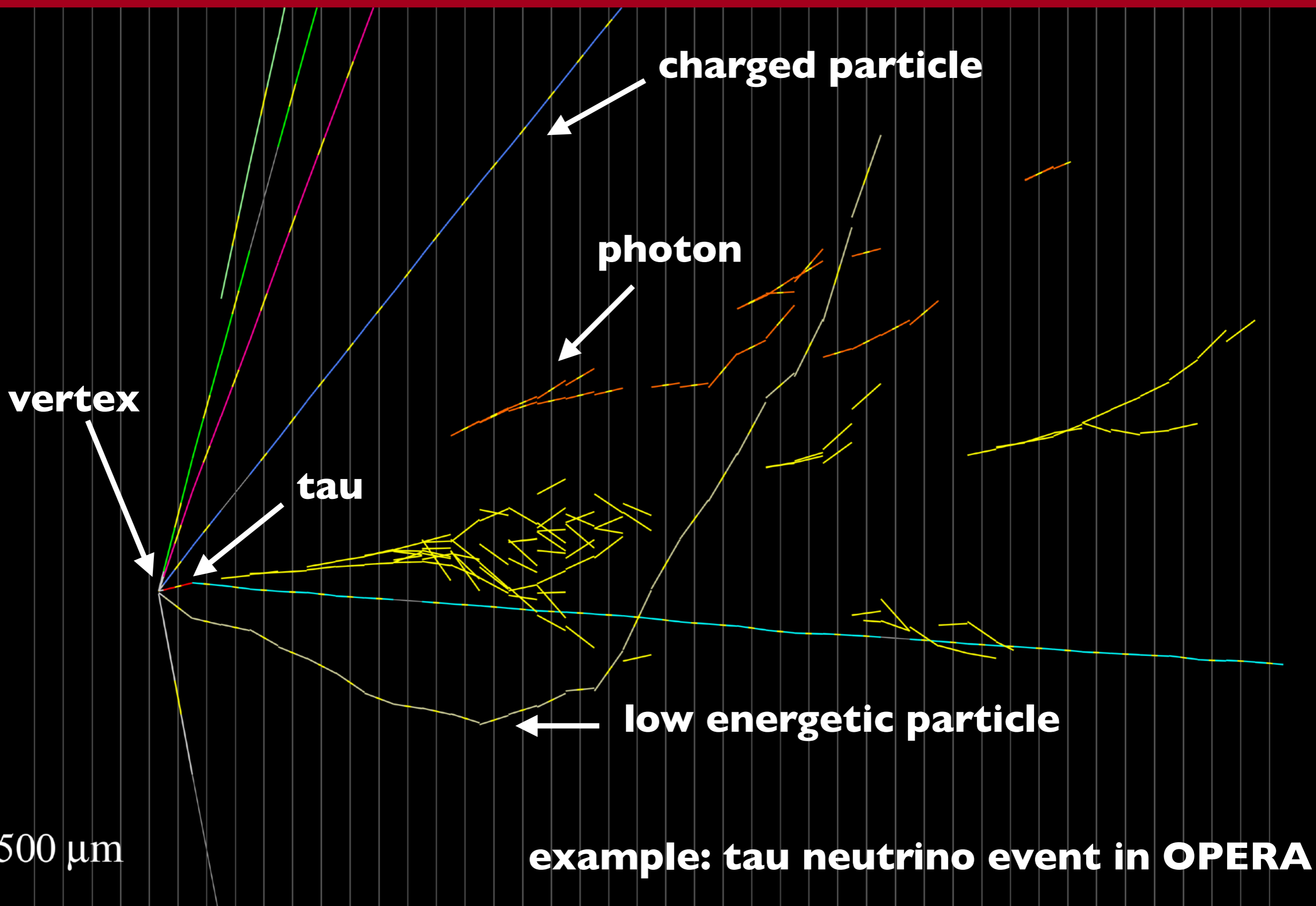
\* 3D tracking devices with 50 nm spatial precision

\* sensitive to neutrino interactions

\* allows particle identification

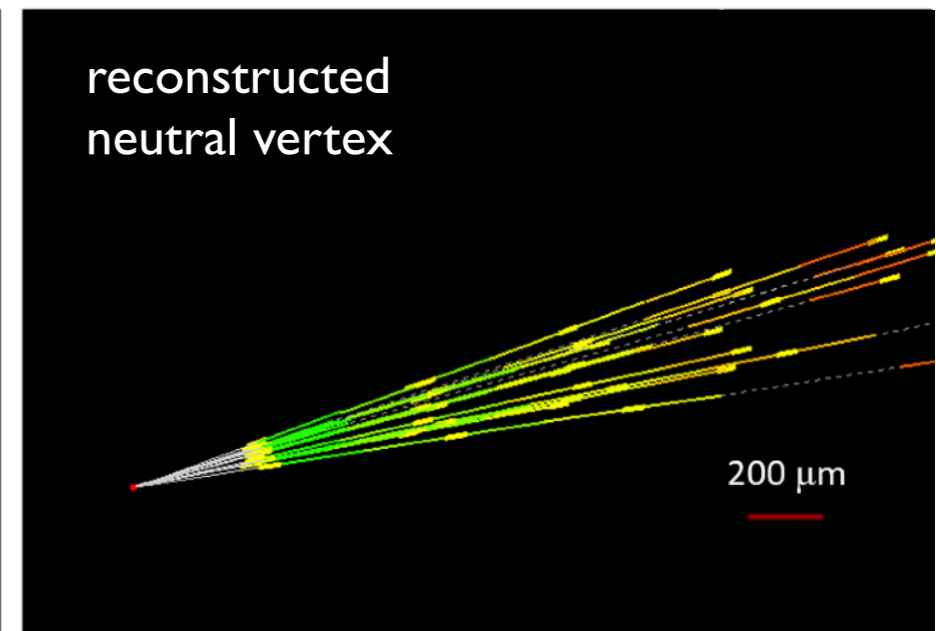
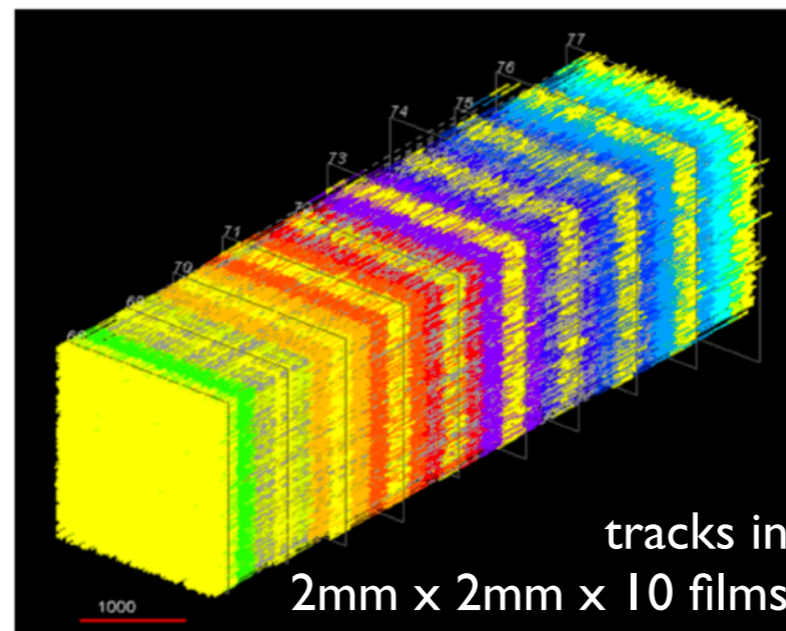
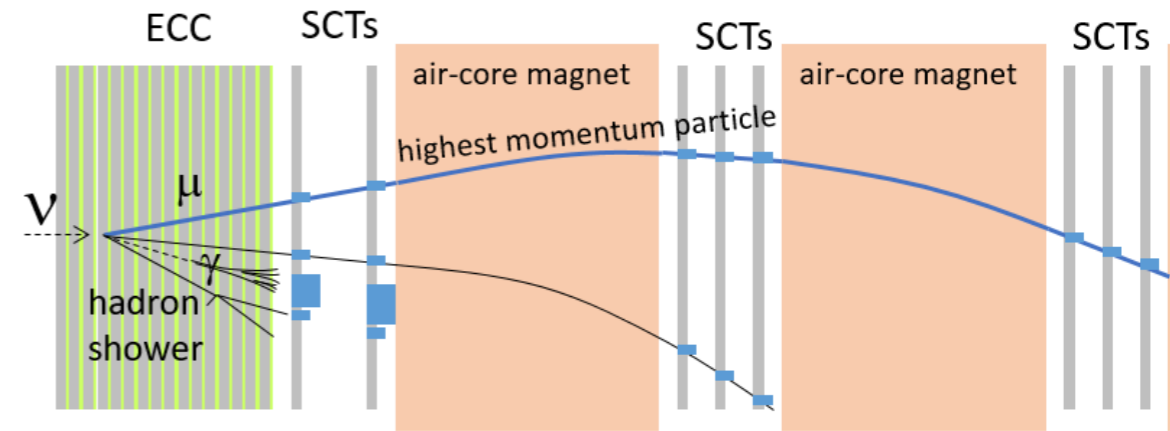


# FASER $\nu$ Detector

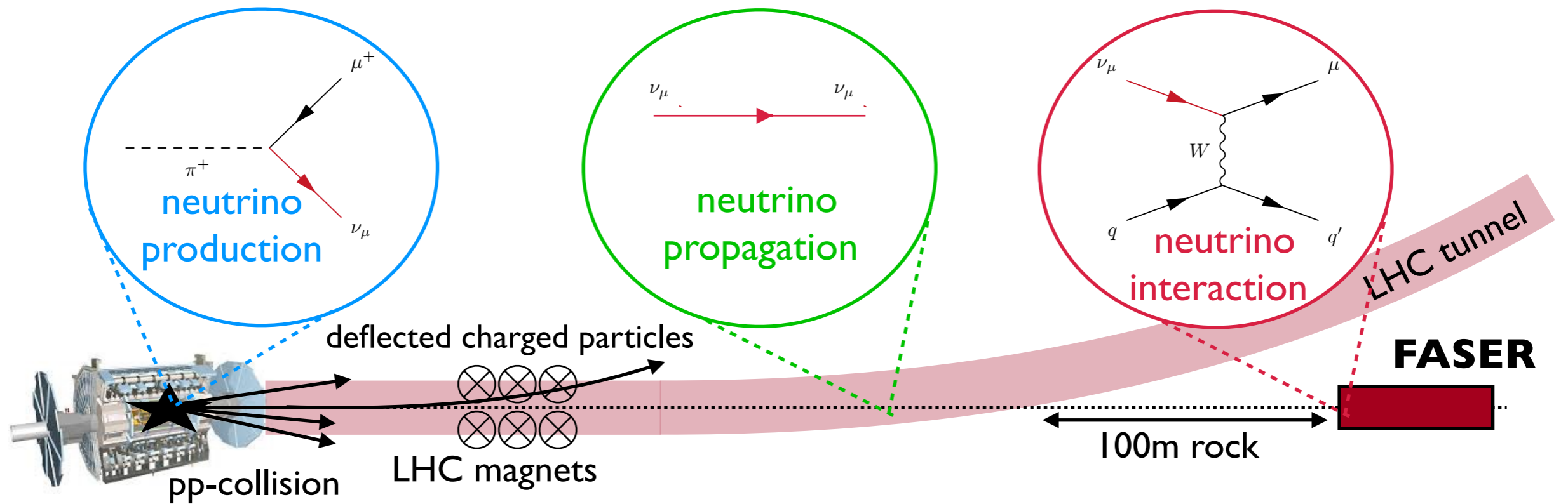


# FASER $\nu$ Detector

- detector performance has been studied
  - \* flavor identification
  - \* vertex finding efficiency:  $\sim 80\%$
  - \* energy resolution:  $\sim 30\%$
- global reconstruction with the FASER detector
  - \* distinguish neutrino / anti-neutrino
  - \* improve neutrino energy reconstruction
  - \* background rejection
- pilot detector data is currently analyzed
  - \* 30 kg detector was installed in T118, 12.5 fb $^{-1}$  of data collected 2018
  - \* goal: first neutrino detection at the LHC



# FASER $\nu$ Physics Potential



In the following, I will present some ideas.

Most of them were not investigated in detail yet.



# FASERν Physics Potential: Neutrino Physics

## - Neutrino cross section at TeV energies

- \* unexplored energy range for all three flavors
- \* charged and neutral current interactions
  - sensitivity to non-standard interactions ?

## - Tau neutrino physics

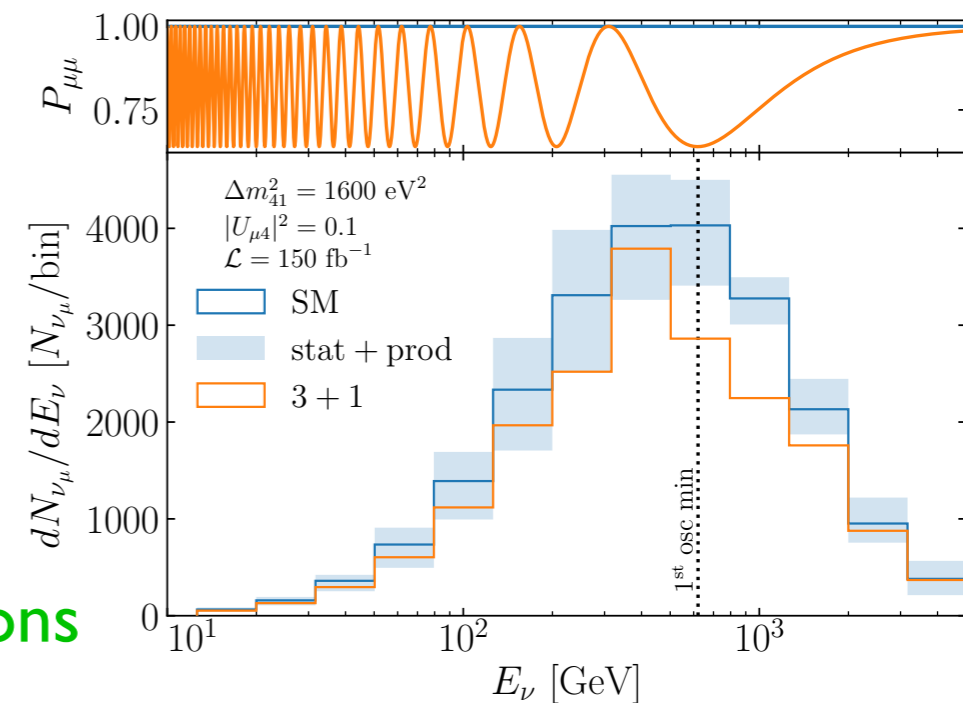
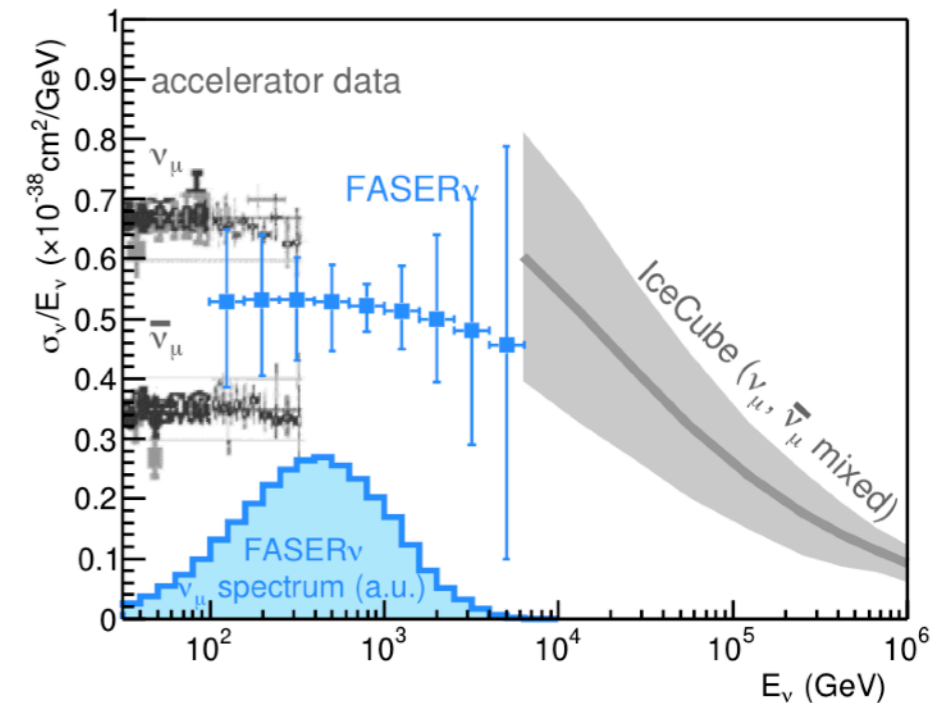
- \* so far only ~20 tau neutrino interactions identified
- \* thousands of tau neutrino events possible at HL-LHC
- \* precision studies of tau neutrino properties
  - tau neutrino cross section
  - tau neutrino magnetic moment ?
  - test lepton flavor universality via  $\nu_\tau c \rightarrow \tau b$  ?

## - Neutrino interaction event shapes

- \* FASERν will record topology/kinematics of interaction
  - validation/tuning of neutrino event generators ?

## - Neutrino oscillations

- \* no neutrino oscillations expected in SM
- \* sterile neutrinos with mass  $\sim 40\text{eV}$  can cause oscillations
- \* use FASERν as short-baseline neutrino experiment



# FASERv Physics Potential: BSM Physics

## - BSM tau-neutrino production

- \* tau neutrino flux small in SM
- \* new light weakly coupled gauge bosons could decay into tau neutrinos: [arXiv:2005.03594](https://arxiv.org/abs/2005.03594)

## - DM scattering

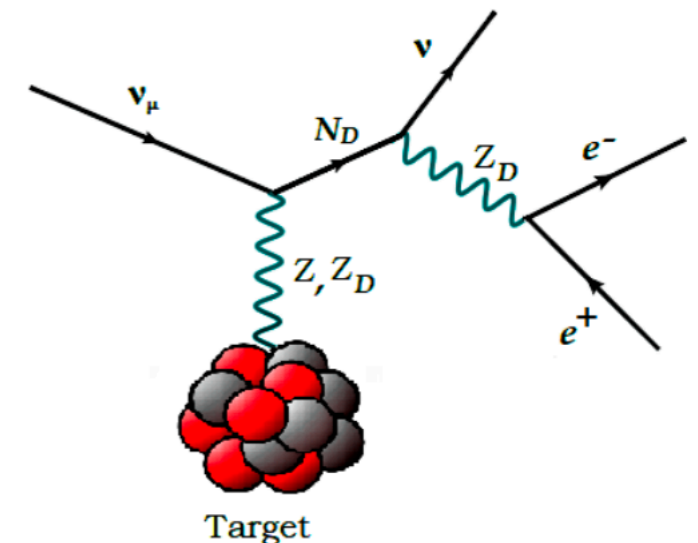
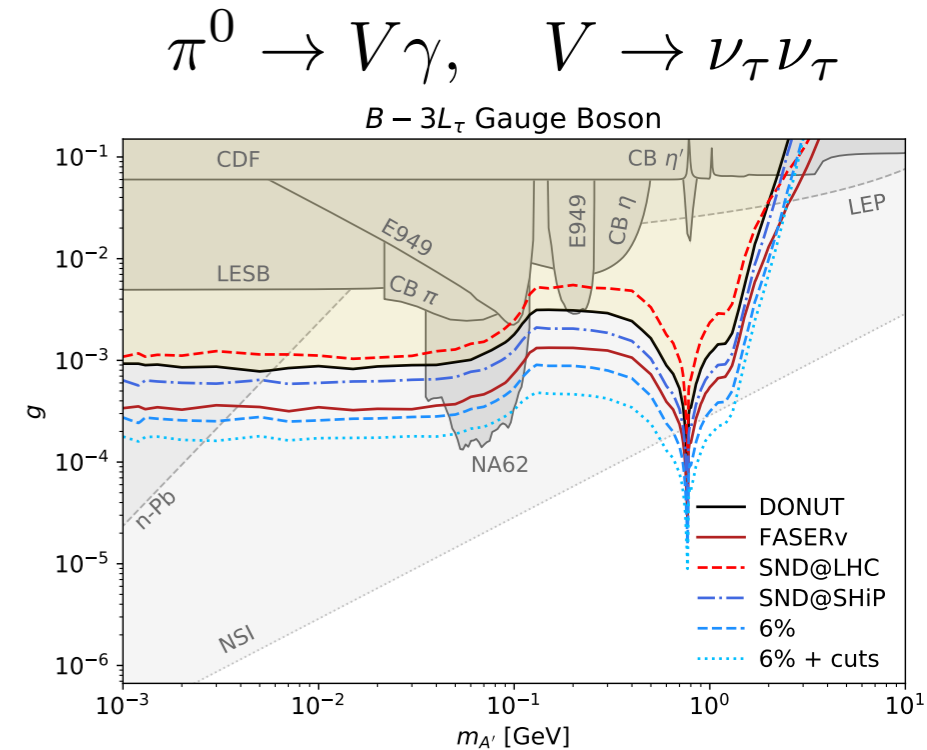
- \* if DM is light, LHC can produce an energetic and collimated DM beam towards FASER
- can we probe DM via scattering in FASERv ?
- \* example:  $\pi \rightarrow V\gamma$  with  $V \rightarrow XX$

## - FASERv as muon beam dump

- \* about  $10^9$  muons will pass FASERv during LHC Run3
- can we use FASERv as muon beam dump experiment ?

## - Probe of neutrino-philic new physics

- \* many models of new physics predict new particles that can interact with neutrinos
- can we use FASERv to probe such models

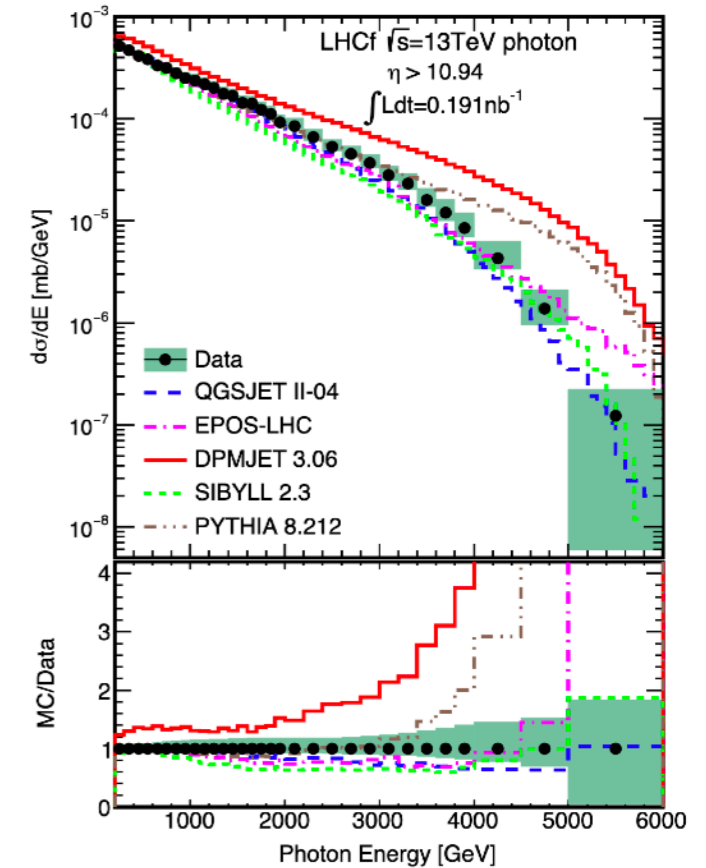


from [arXiv:1807.09877](https://arxiv.org/abs/1807.09877)

# FASER $\nu$ Physics Potential: QCD

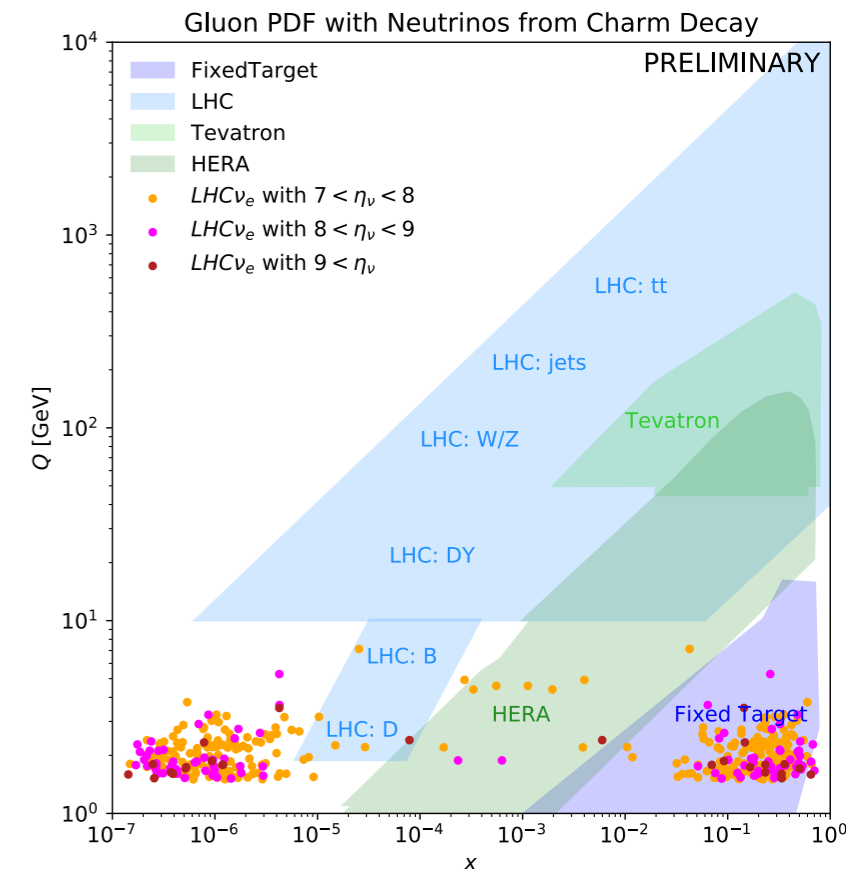
## - Constraining forward particle production

- \* forward particle production is poorly constrained
- \* FASER $\nu$  neutrinos flux measurements will provide complimentary constraints
- \* light hadrons described by non-perturbative physics
  - validate/improve these hadronic interaction models ?
- \* heavy hadrons can be described by perturbative QCD
  - constrain low-x gluon PDF ?
  - probe intrinsic charm ?



## - PDFs via Neutrino scattering

- \* DIS neutrino scattering is sensitive to PDFs
  - nuclear PDFs at a variety of targets ?
  - $\nu s \rightarrow l c$  as probe of the strange PDF ?



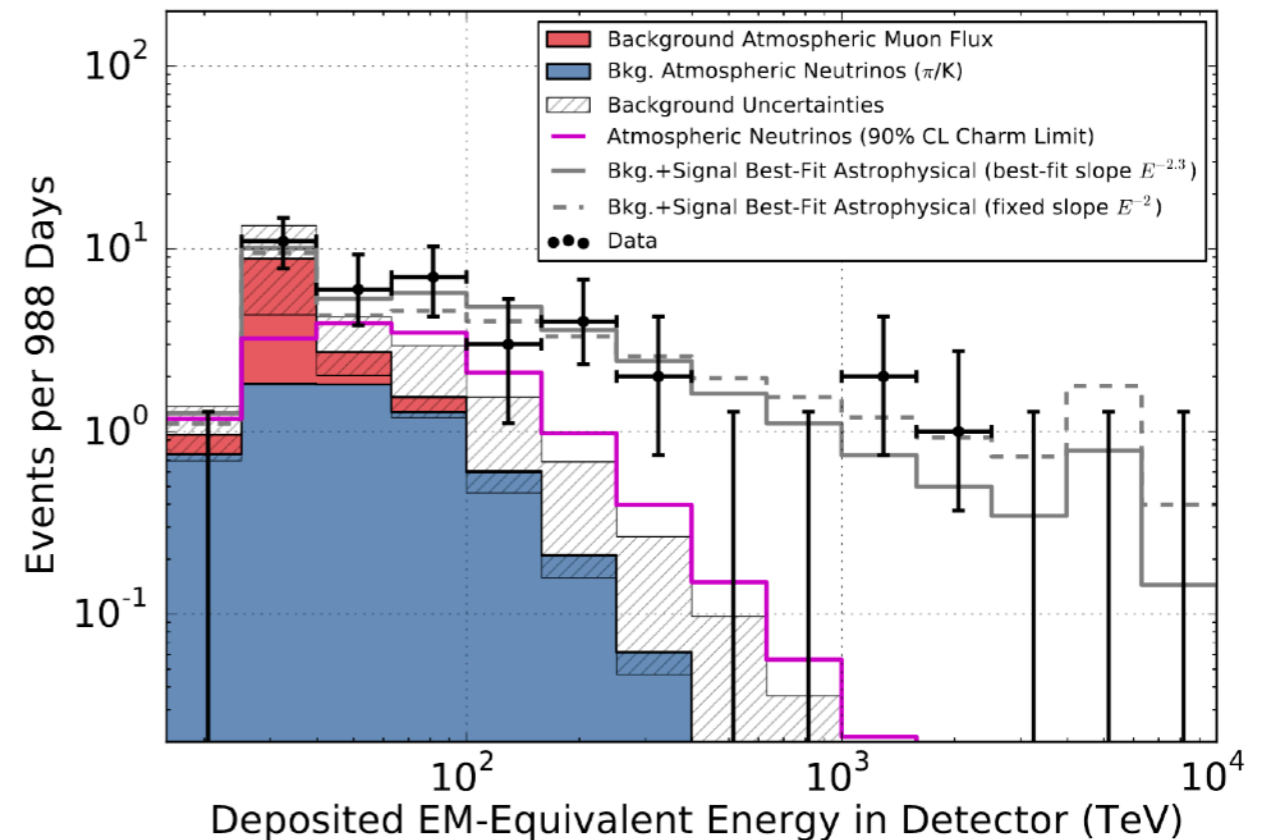
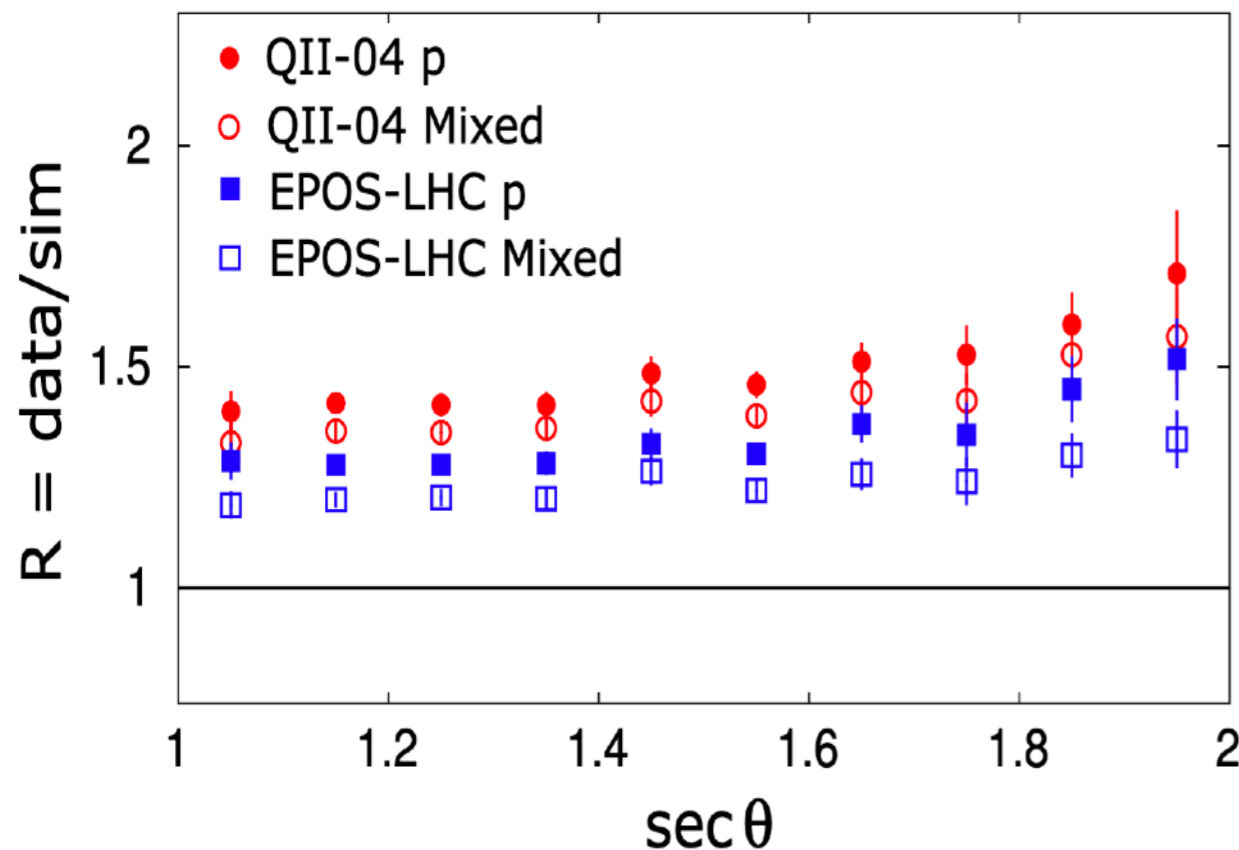
# FASERv Physics Potential: Comics

## - Muon problem in CR physics

Cosmic Ray experiments have reported an excess in the number of muons over expectations computed using extrapolations of hadronic interaction models tuned to LHC data at the few  $\sigma$  level. See [1610.08509](#)

## - Prompt Atmospheric Neutrinos

Prompt Atmospheric Neutrinos from charm decay are the dominant background for IceCube's comic neutrinos, and we currently have no data and no theory for this process. See [1405.5303](#)

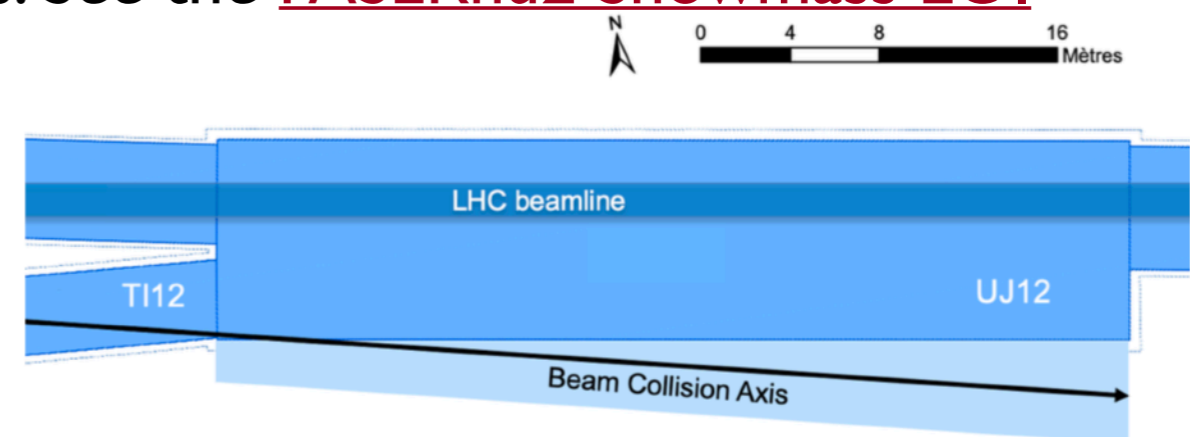




## Summary and Outlook

# Planning for the future

- FASER and FASERv are approved, funded and under construction. They will take data during Run3 of the LHC (2021-2024), but also pave the way for a high energy forward search and neutrino program, opening up many new opportunities for [neutrino physics](#), [new physics searches](#) and [QCD](#), significantly extending the LHC's physics program.
- As part of the Snowmass 2021 community study, we are planning to investigate the full physics potential of FASER and FASERv during the HL-LHC era.
  1. We propose an upgraded FASER2 detector with enhanced sensitivity for LLP searches. See the [FASER2 Snowmass LOI](#)
  2. We propose an enlarged neutrino detector with roughly ten times the mass of FASERnu operating at the HL LHC. Such a detector would collect roughly  $\sim 10^5 \nu_e$ ,  $\sim 10^6 \nu_\mu$  and  $\sim 10^3 \nu_\tau$  at TeV energies. See the [FASERnu2 Snowmass LOI](#)
  3. We proposal to enlarge an existing cavern and create a Forward Physics Facility to house a variety of experiments. See the [Forward Physics Facility Snowmass LOI](#)
- We would like to invite the HEP community (so you) to help us explore the physics potential of this program.



# Summary and Outlook

## FASER

- newest experiment at the LHC
- quick, small and inexpensive
- funded and approved

## Envisioned Timeline

- build/install FASER in LS2 (2019-20)
- take data during Run 3 (2021-23, 150 fb<sup>-1</sup>)
- upgrade to FASER 2 in LS3 for HL-LHC

## Physics:

- search for light long-lived particles at the LHC
- neutrino measurements at TeV energies
- and many more unexplored opportunities ...  
→ help us explore them

For more information, see our website: <https://faser.web.cern.ch/>

SIMONS  
FOUNDATION



Many thanks to the Heising-Simons Foundation, the Simons Foundation, and to CERN for invaluable support

**We look forward to feedback and suggestions**

# Backup



# FASER Timeline



**09/2017** - Original FASER idea paper [arXiv:1708.09389](https://arxiv.org/abs/1708.09389)

**spring 2018** - FASER collaboration forms

**07/2018** - Letter of Intent [arXiv:1811.10243](https://arxiv.org/abs/1811.10243)

**11/2018** - Technical Proposal [arXiv:1812.09139](https://arxiv.org/abs/1812.09139)

**11/2018** - LLP Physics Potential [arXiv:1811.12522](https://arxiv.org/abs/1811.12522)

**12/2018** - Funding

**03/2019** - Approval by CERN

**08/2019** - Proposal of FASERv  
[arXiv:1908.02310](https://arxiv.org/abs/1908.02310)

**11/2019** - FASERv TP  
[arXiv:2001.03073](https://arxiv.org/abs/2001.03073)

65 collaborators  
18 institutions  
8 countries



Detector Construction and Integration - **2019-2020**

Collecting Data - **2021-2023**

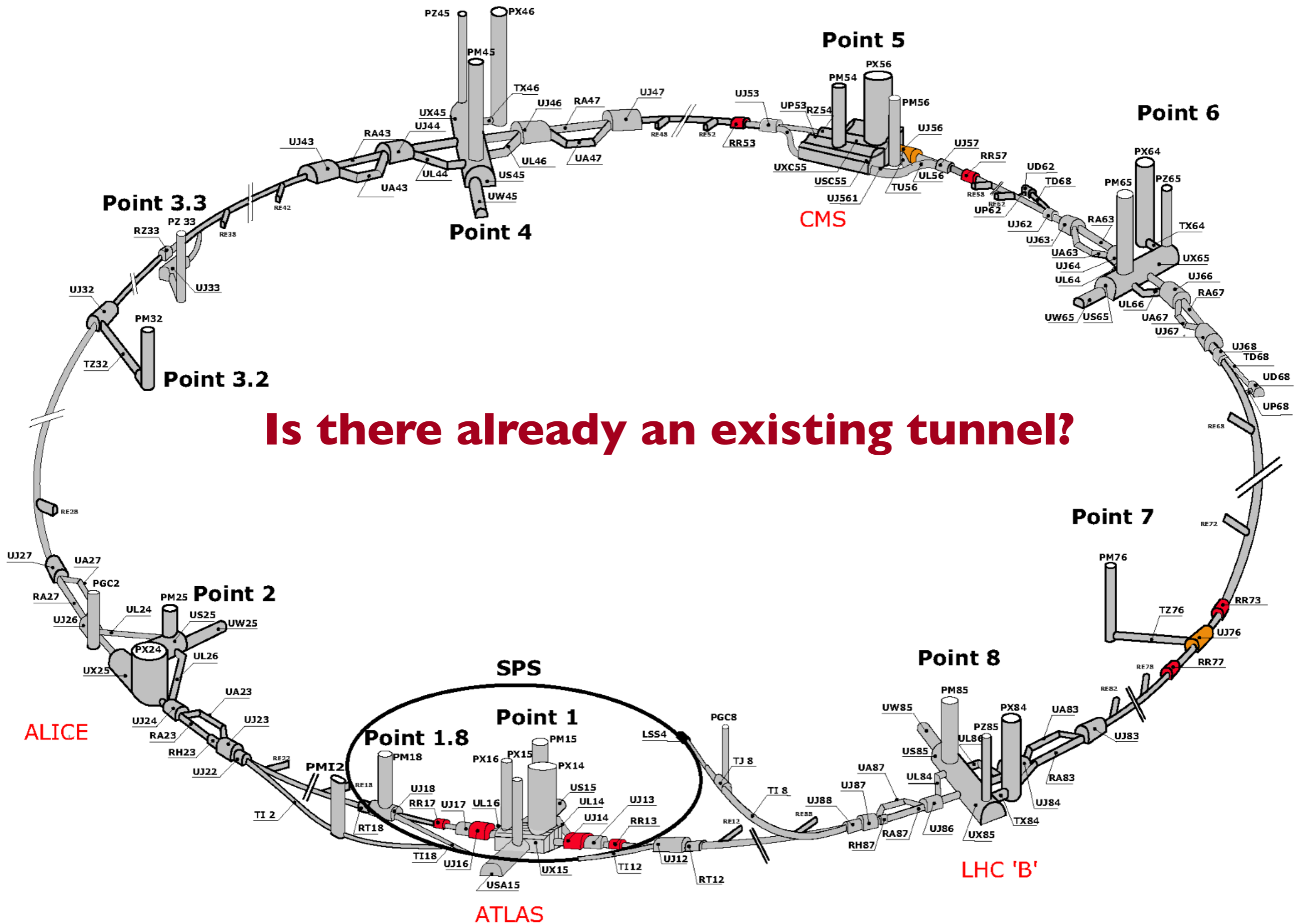
FASER 2 Upgrade - **HL-LHC era**

# Acknowledgements

The FASER Collaboration gratefully acknowledges the contributions of many people.

We are grateful to the ATLAS SCT project and the LHCb Calorimeter project for letting us use spare modules as part of the FASER experiment. In addition, FASER gratefully acknowledges invaluable assistance from many people, including the CERN Physics Beyond Colliders study group; the LHC Tunnel Region Experiment (TREX) working group; Rhodri Jones, James Storey, Swann Levasseur, Christos Zamantzas, Tom Levens, Enrico Bravin (beam instrumentation); Dominique Missiaen, Pierre Valentin, Tobias Dobers (survey); Jonathan Gall, John Osborne (civil engineering); Caterina Bertone, Serge Pelletier, Frederic Delsaux (transport); Francesco Cerutti, Marta Sabaté-Gilarte, Andrea Tsinganis (FLUKA simulation and background characterization); Pierre Thonet, Attilio Milanese, Davide Tommasini, Luca Bottura (magnets); Burkhard Schmitt, Christian Joram, Raphael Dumps, Sune Jacobsen (scintillators); Dave Robinson, Steve McMahon (ATLAS SCT); Yuri Guz (LHCb calorimeters); Salvatore Danzeca (Radiation Monitoring); Stephane Fartoukh, Jorg Wenninger (LHC optics), Michaela Schaumann (LHC vibrations); Marzia Bernardini, Anne-Laure Perrot, Katy Foraz, Thomas Otto, Markus Brugger (LHC access and schedule); Simon Marsh, Marco Andreini, Olga Beltramello (safety); Stephen Wotton, Floris Keizer (SCT QA system and SCT readout); Liam Dougherty (integration); Yannic Body, Olivier Crespo-Lopez (cooling/ventilation); Yann Maurer (power); Marc Collignon, Mohssen Souayah (networking); Gianluca Canale, Jeremy Blanc, Maria Papamichali (readout signals); Bernd Panzer-Steindel (computing infrastructure); and Mike Lamont, Fido Dittus, Andreas Hoecker, Andy Lankford, Ludovico Pontecorvo, Michel Raymond, Christoph Rembser, Stefan Schlenker (useful discussions).

# FASER Location

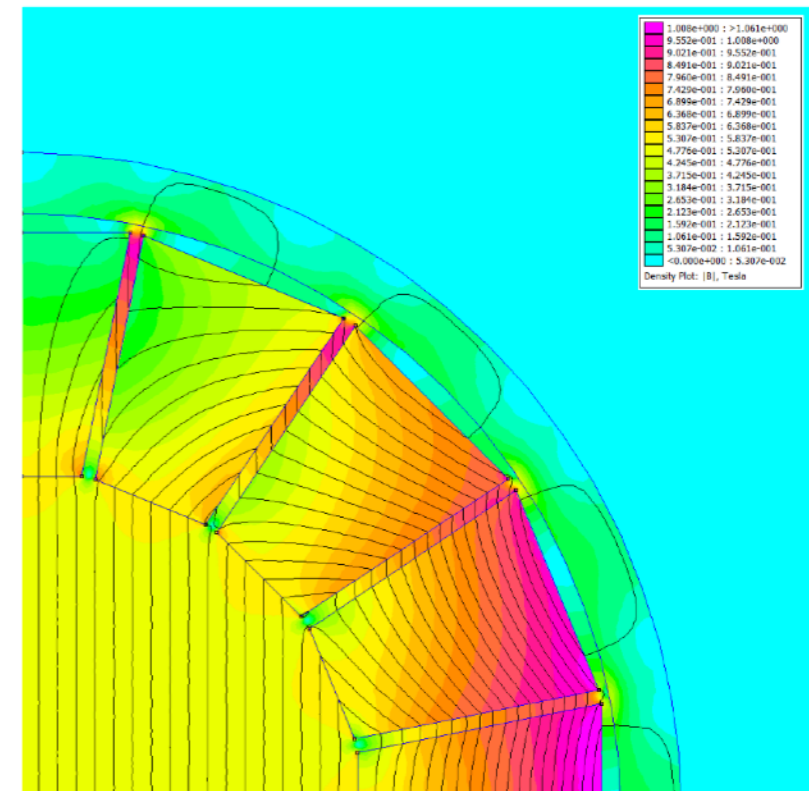
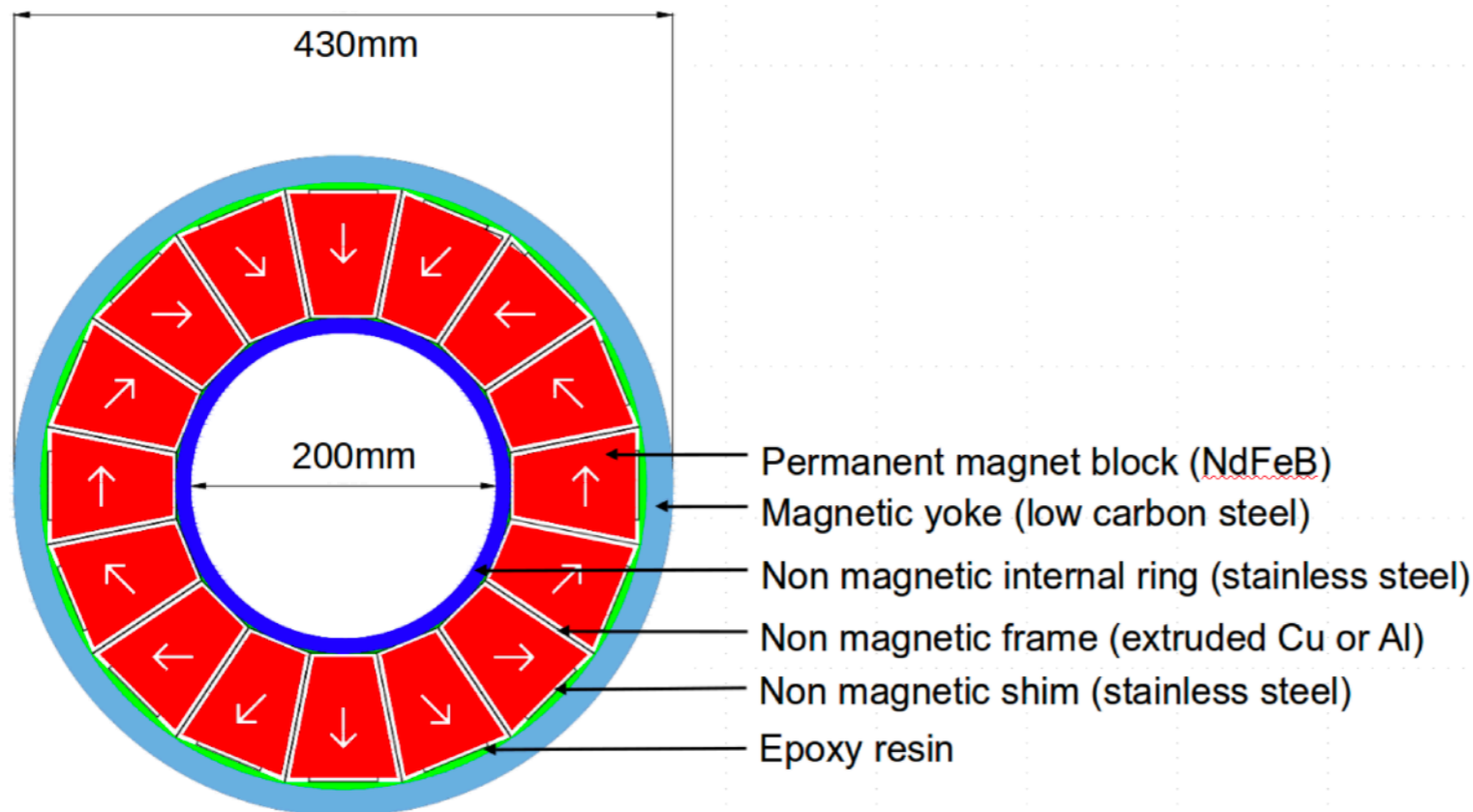
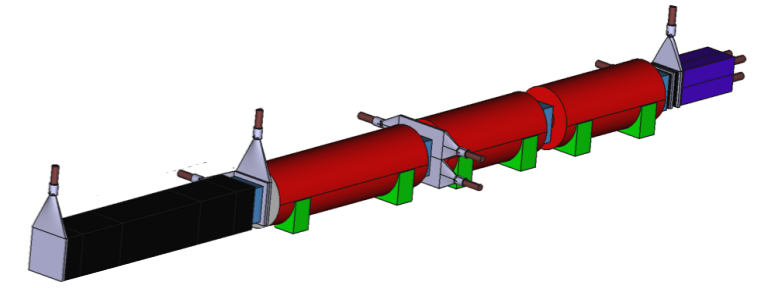


**Is there already an existing tunnel?**

# FASER Detector

## FASER Magnet

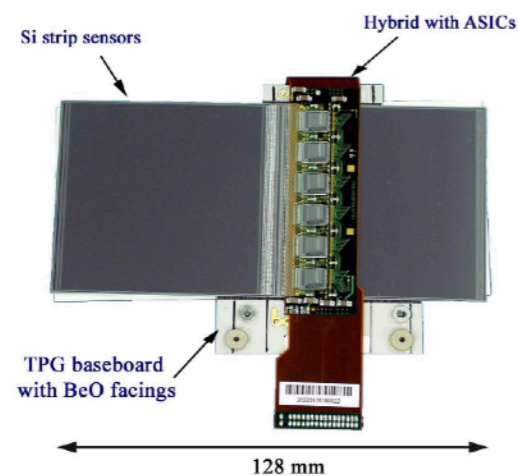
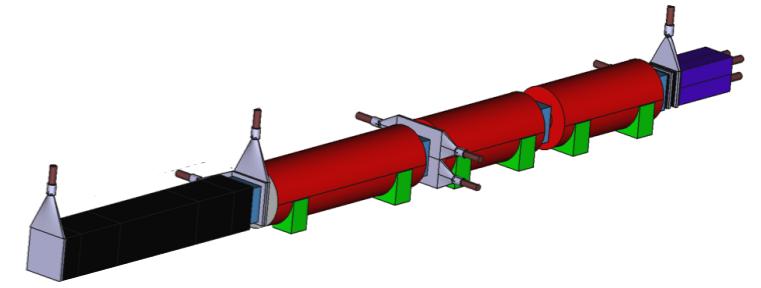
- 0.6T permanent dipole magnets
- Halbach array design
  - LOS to passes through the magnet center
  - minimum digging to the floor in T112
  - minimized needed services (power, cooling etc..)
- to be constructed by the CERN magnet group



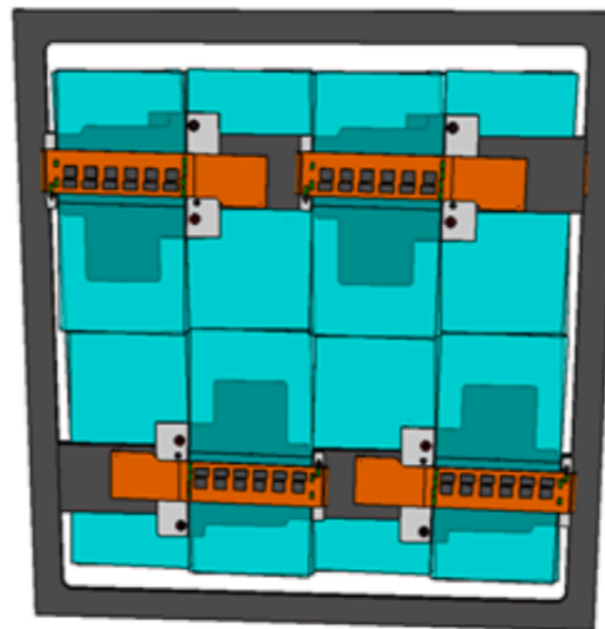
# FASER Detector

## FASER Tracker

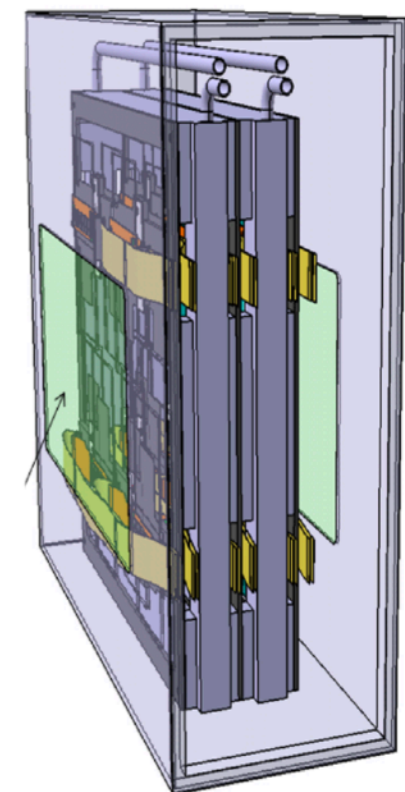
- 3 tracking stations, each with 3 tracking layers
- double sided silicon micro-strip detectors
- ATLAS SCT spare modules will be used
  - 80 $\mu$ m strip pitch, 40mrad stereo angle
  - many thanks to the ATLAS SCT collaboration!
  - 72 SCT modules for the full tracker



SCT module



Tracking layer

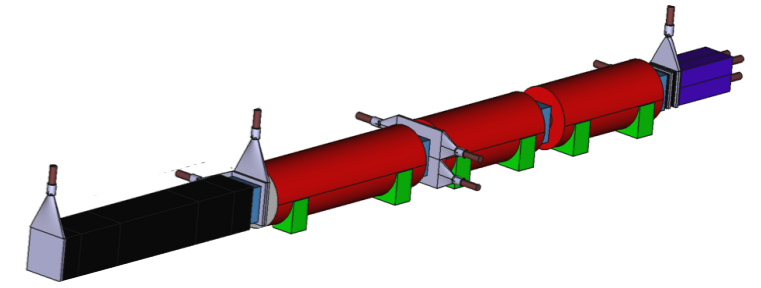


Tracking station

# FASER Detector

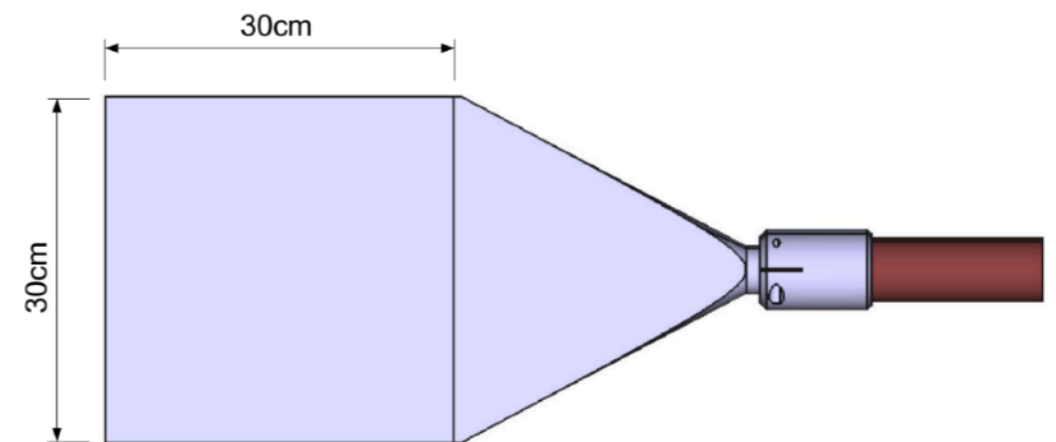
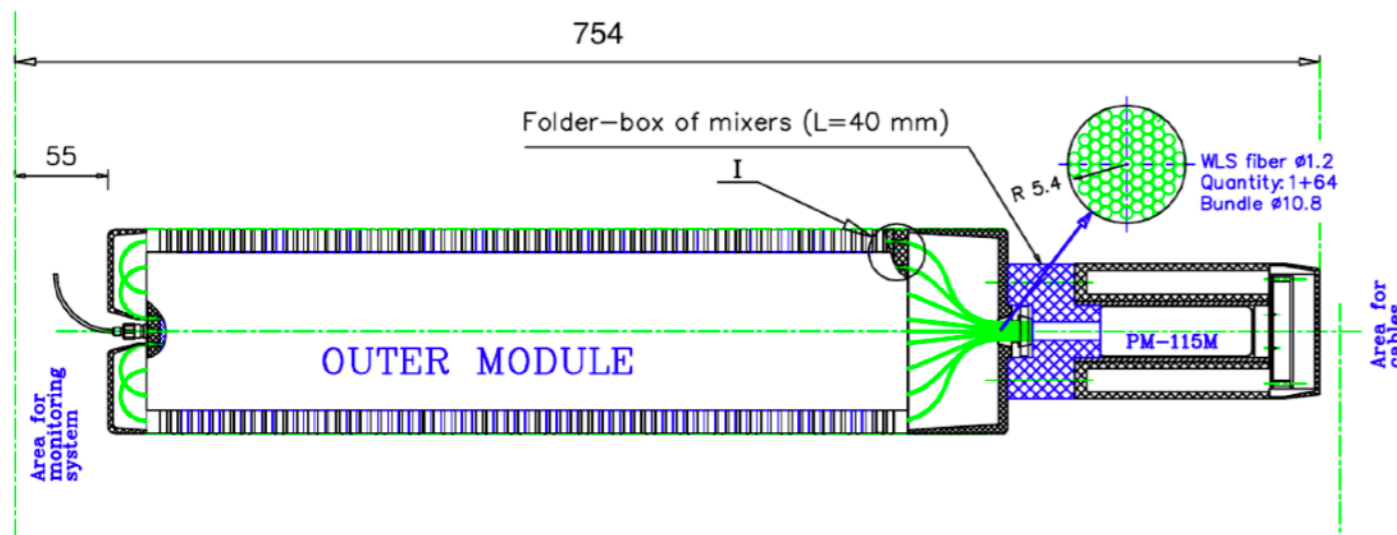
## FASER ECAL

- EM energy measuring / triggering / electron/photon identification
- FASER will use spare LHCb outer ECAL modules
  - $\sim 1\%$  energy resolution for 1 TeV electrons
  - Many thanks for LHCb for allowing us to use these!



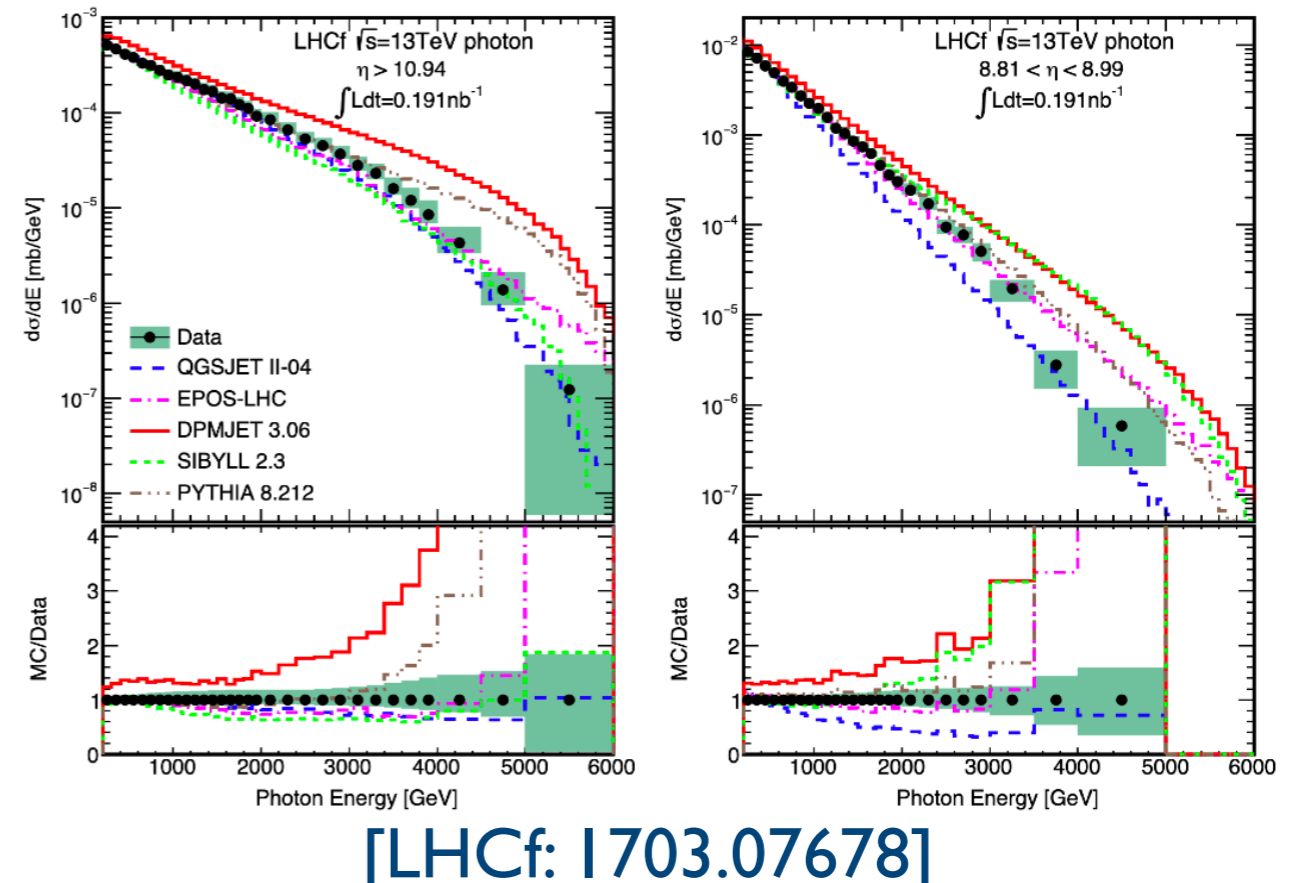
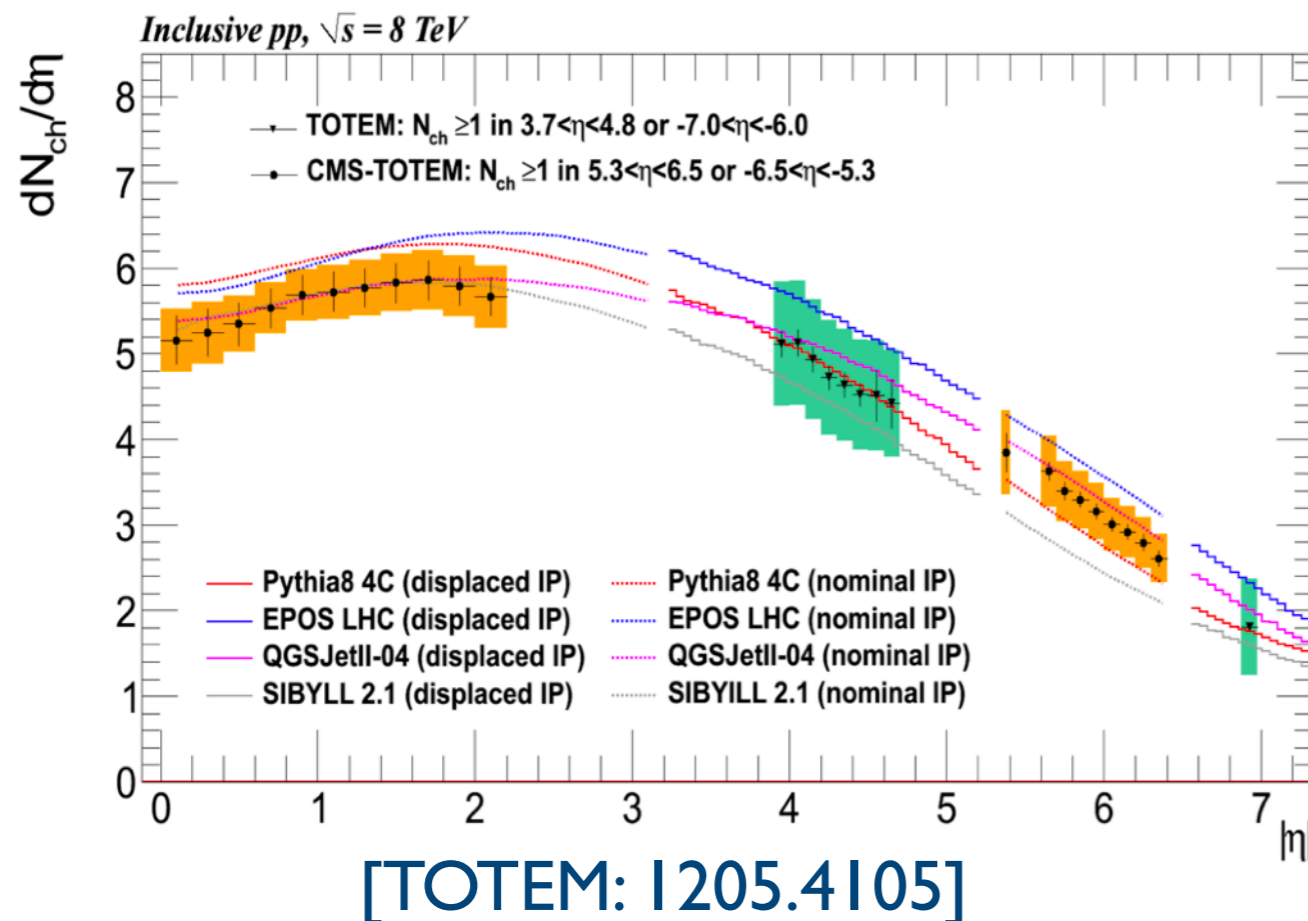
## FASER Scintillators

- vetoing charged particles entering the decay volume / triggering
- to be produced at CERN scintillator lab



# FASERv Physics

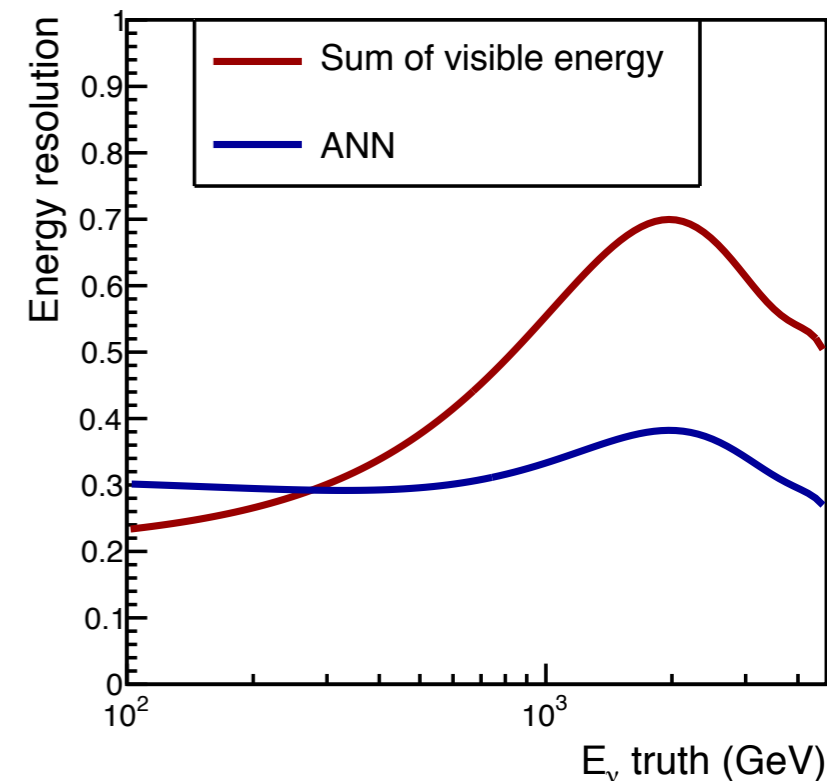
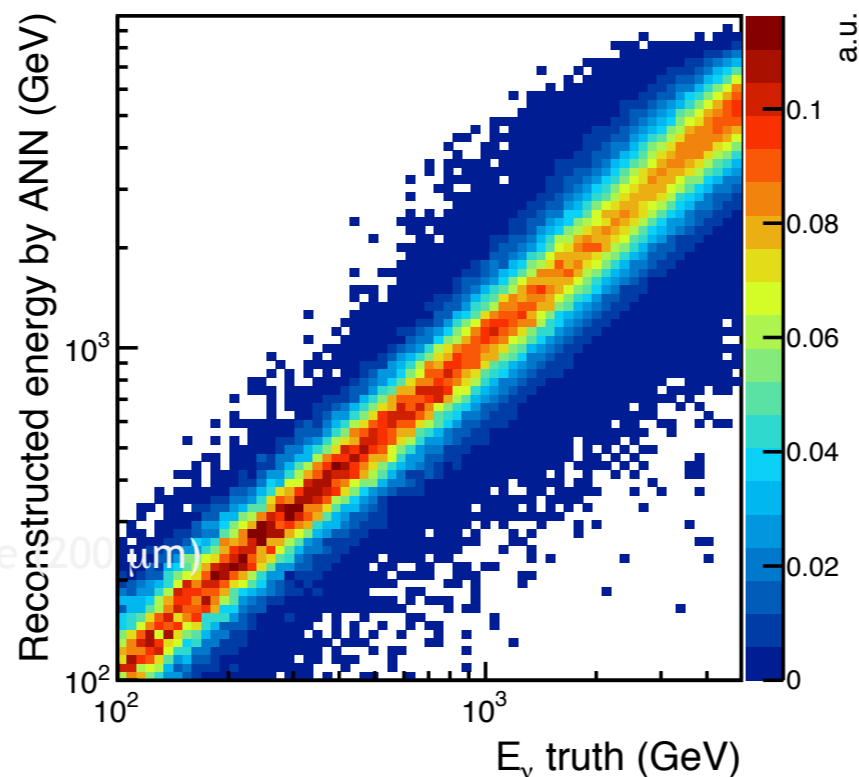
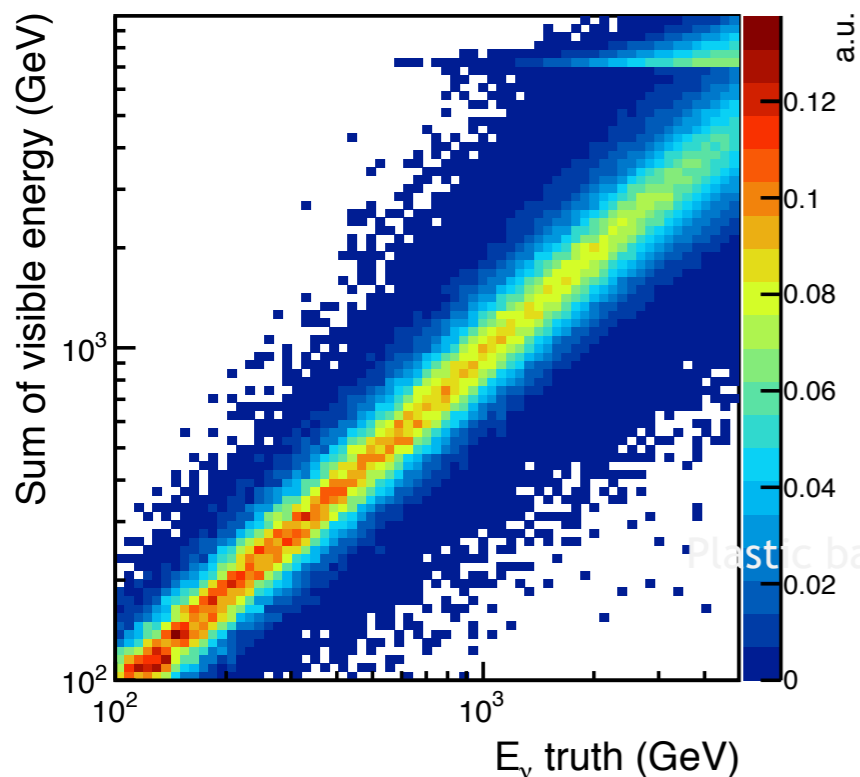
- cross section measurements limited by neutrino flux uncertainty
    - \* we need to **quantify** and **reduce** these uncertainties
  - forward particle production not described by pQCD, but soft physics
    - \* use hadronic interaction models
  - simulators based on sophisticated modeling of microscopic physics
    - \* phenomenological parameters need to be tuned
    - \* include tuning uncertainties (similar to PDFs)
- **develop dedicated forward physics tune using forward data**



# FASER $\nu$ Detector

- neutrino energy reconstruction
  - \* use topological + kinematic observables
  - \* train NN to estimate neutrino energy
  - \* 30% energy resolution seems achievable

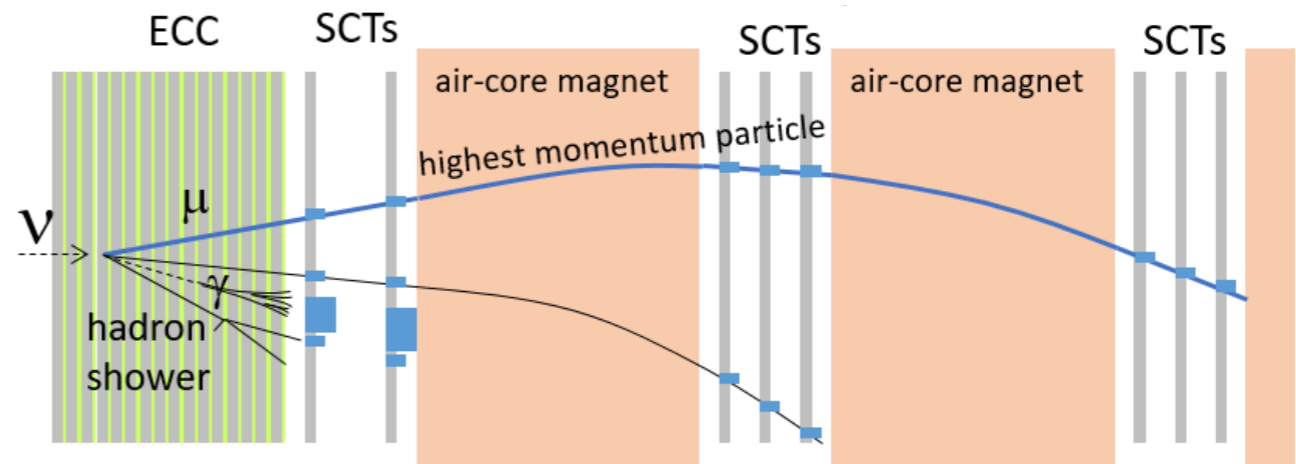
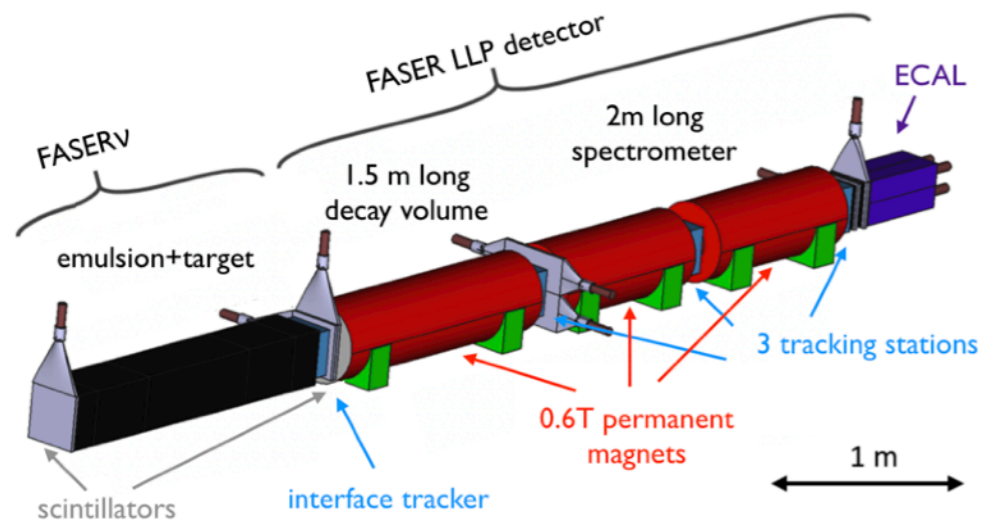
Topological Variables		related to
$n_{\text{tr}}$	charged tracks multiplicity	$E_{\text{had}}$
$n_{\gamma}$	photon multiplicity	$E_{\text{had}}$
$ 1/\theta_{\ell} $	lepton angle	$E_{\ell}$
$\sum  1/\theta_{\text{had}} $	sum of inverse hadron track angles	$E_{\text{had}}$
$1/\theta_{\text{median}}$	inverse of median of all track angles	$E_{\text{had}}, E_{\ell}$
Kinematical Variables		
$p_{\ell}^{\text{MCS}}$	lepton momentum from MCS	$E_{\ell}$
$\sum p_{\text{had}}^{\text{MCS}}$	charged hadron momenta from MCS	$E_{\text{had}}$
$\sum E_{\gamma}$	energy in photon showers	$E_{\text{had}}$



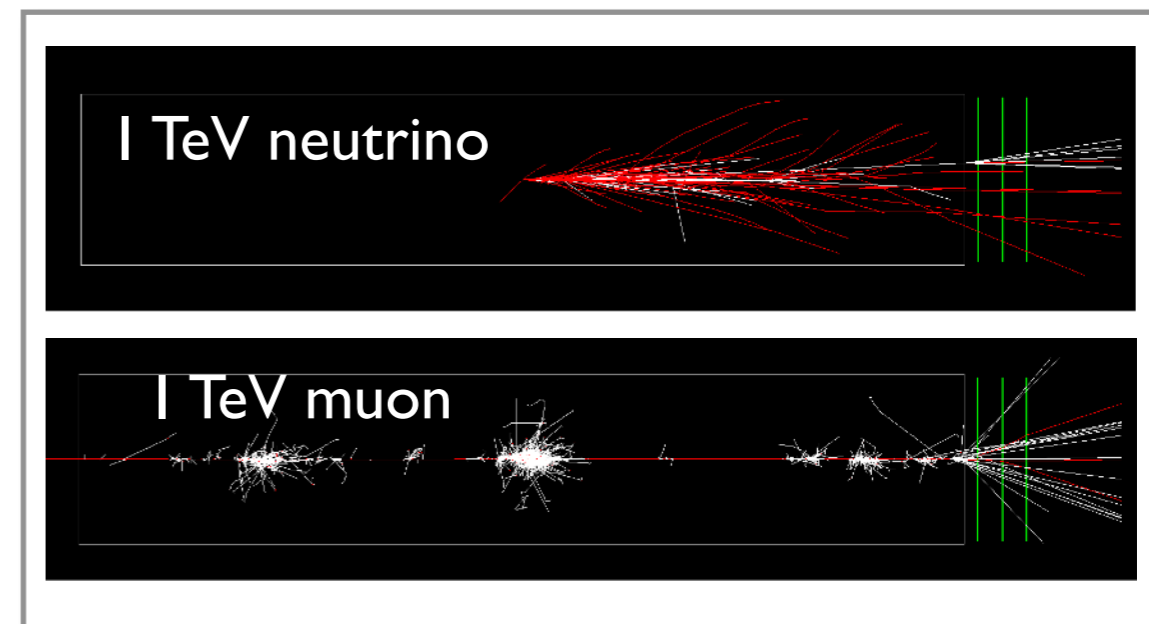


# FASER $\nu$ Detector

- possibility for future: global reconstruction with the FASER detector
  - \* interface FASER $\nu$  to the FASER spectrometer
  - \* requires additional tracking layer to allow matching

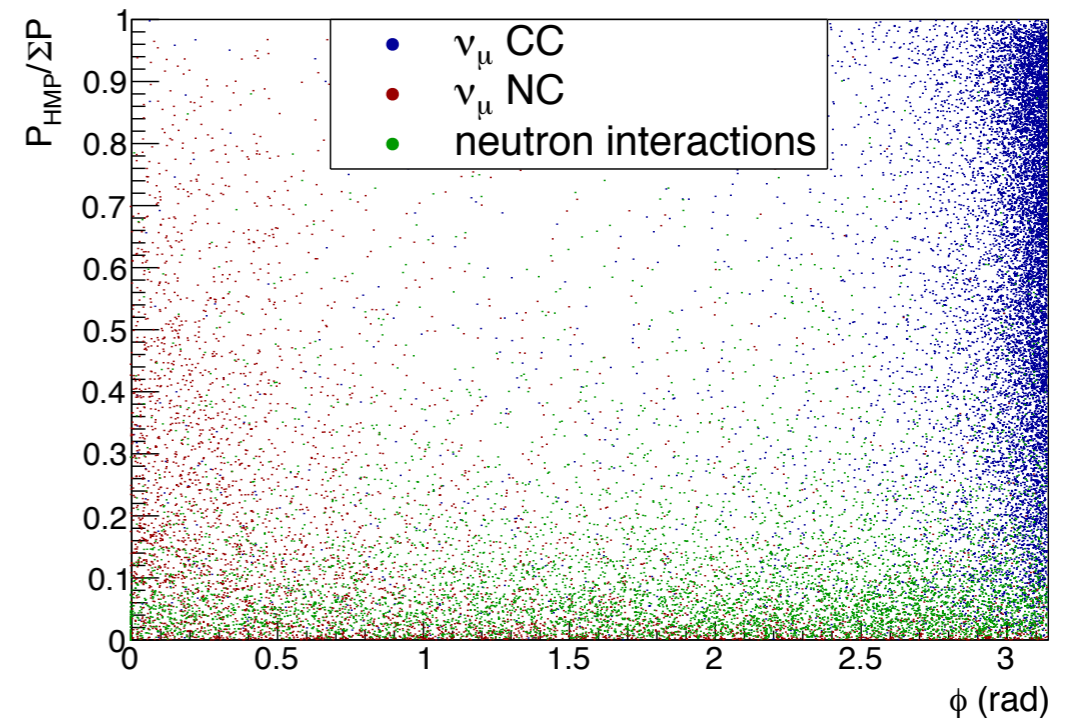
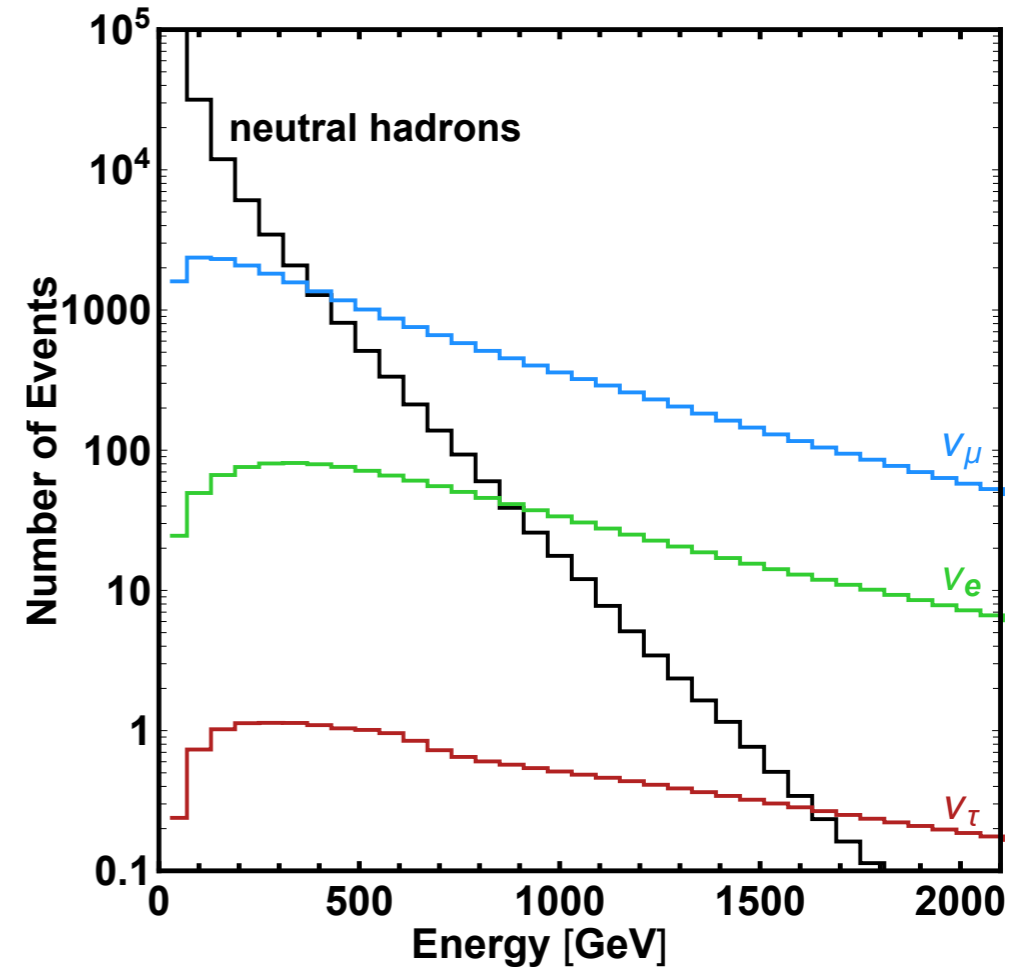


- muon charge identification
  - distinguish neutrino / anti-neutrino
- momentum of charged tracks
  - improve neutrino energy reconstruction
- timestamp events and identify additional activity (muon)
  - background rejection



# FASER $\nu$ Backgrounds

- neutrino interaction signal is striking
  - \* neutral vertex identified with  $>5$  tracks
  - \* points back to IP
  - \* high energy event
  - \* charged lepton identified
- neutral hadrons could mimic neutrino signal
  - \* simulated with FLUKA
  - \* rejection based on energy, lepton ID and kinematics
- lepton identification
  - \*  $\nu_\mu$ : highest momentum particle without interactions
  - \*  $\nu_e$ : single-electron initiated EM shower
  - \*  $\nu_\tau$ : kink from displaced  $\tau$ -decay vertex

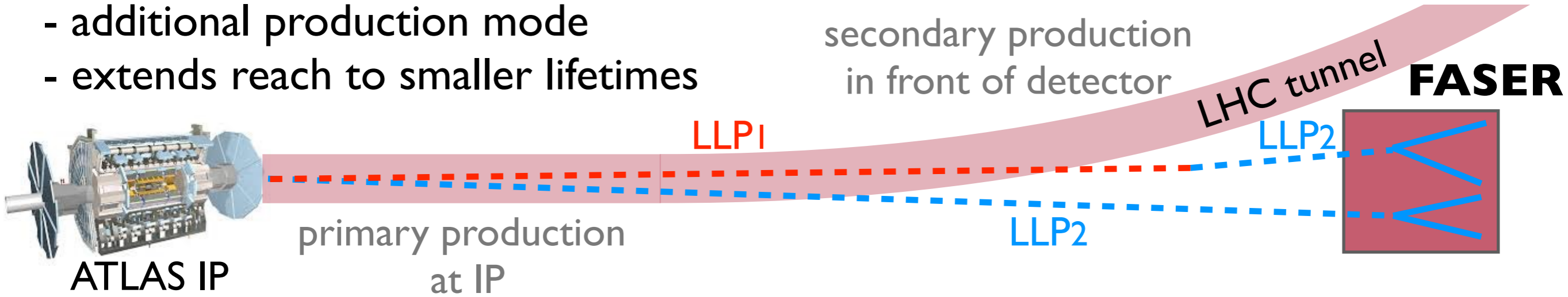


# Long Lived Particles at FASER

## Secondary LLP Production

[K. Jodłowski, FK, L. Roszkowski, S. Trojanowski: 1810.01879]

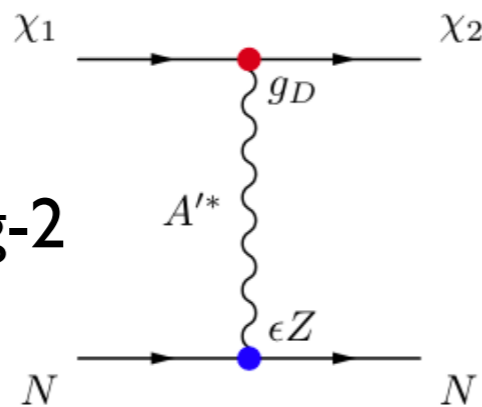
- additional production mode
- extends reach to smaller lifetimes



## Example: inelastic DM

- $A'$  couples diagonally to  $X_1 + X_2$
- $X_1 = \text{LLP}_1$ ,  $X_2 = \text{LLP}_2$
- primary production  $\pi^0 \rightarrow \gamma X_1 X_2$
- secondary production:  $X_1 N \rightarrow X_2 N$
- decay:  $X_2 \rightarrow X_1 e e$

- enhanced reach
- relevant for DM &  $g-2$



iDM:  $m_{\chi_1} : m_{\chi_2} : m_{A'} = 1 : 2.9 : 4$ ,  $\alpha_D = 0.1$

